

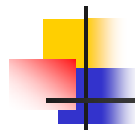
A Treatise on Rotation of Celestial Bodies. Part I



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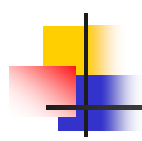
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0. Summary

- Motion of Finite Body
 - = Orbital Motion + Rotation
- Basics: Rotation of Rigid Body
 - Orientation: Rotation Matrix
 - Eq. of Motion: Angular Momentum Conservation
 - Solution: Free Rotation + Perturbation
- Topics: Non-Rigidity, Relativity
- Appl.: Earth, Moon, ...



Various Rotation

- Earth: $A \sim B < C$, $C - A \ll C$
 - Almost Uniform Rotation: UT1
 - Forced Motion of Rot. Axis: Precession, ...
 - Non-Rigidity: Polar Motion, ...
 - Similar Case: Mars
- Moon: $A < B < C$, $B - A < C - A \ll C$
 - Strong Spin-Orbit Coupling
 - Similar Case: Galilean Satellites



Various Rotation (2)

- Asteroids
 - Free Rotation + Small Sun's Torque
 - Large Non-Sphericity: $A < B < C$
 - Sometimes Prolate: Ex.) Ida, $A < B \sim C$
- Artificial Satellites
 - Arbitrary Shape = Arbitrary (A,B,C)
 - Strong Perturbation: Earth's Torque, ...
 - Active Control of Attitude



History

- Hipparchus (~BC150): Precession
- Copernicus (1543): Earth's Axis Motion
- Fabricius (1611): Sun's Rotation
- Huygens (1673): Centrifugal Force
- Newton (1687): Law of Motion, Universal Attraction, Explanation of Precession
- Cassini I (1693): Law of Moon's Rotation



History (2)

- D'Alembert (1743): Prediction of Nutation
- Bradley (1747): Discovery of Nutation
- Euler (1765): Eq. of Rigid Body Rotation, Prediction of Polar Motion
- Landen (1775): Landen's Transf.
- Legendre (1786): Elliptic Integrals



History (3)

- Lagrange (1788): Analytical Dynamics
- Poisson (1809): Poisson Approximation
- Jacobi (1829): Elliptic Functions
- Poincaré (1834): Visualization of Rotation
- Hamilton (1834): Canonical Formulation
- Coriolis (1835): Coriolis Force



History (4)

- Serret (1866): Serret Canon. Variables
- Oppolzer (1880): Oppolzer Term
- Chandler (1891): Discovery of Polar Motion
- 1899~: International Latitude Service
- Kimura (1902): Discovery of Z-Term
- Poincare (1910): Fluid Core Rotation
- Einstein (1915): General Relativity



History (5)

- De Sitter (1917): Geodesic Precession
- Andoyer (1923): Canonical Theory of Rotation
- Oort (1927): Galactic Rotation
- Woolard (1953): Modern Nutation Theory
- Molodensky (1961): Nutation of Elastic Earth
- 1962~: Int'l Polar Motion Service



History (6)

- Hori (1966): Pert. Theory by Lie-Transf.
- 1969~: Lunar Laser Ranging Obs.
- Wako (1970): Clarification of Z-Term
- 1976~: Satellite Laser Ranging Obs.
- Lieske et al. (1977): IAU1976 Precession
- Kinoshita (1977): Rigid Earth Nutation Theory based on Hori's Perturbation Theory



History (7)

- Guinot (1979): Non-Rotating Origin
- 1980: MERIT Campaign
- Sasao et al. (1980): SOS Theory
- Wahr (1981): Nutation Theory of Non-rigid Earth
- Seidelmann (1982): IAU1980 Nutation Theory



History (8)

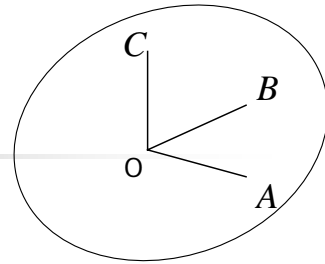
- Eckhardt (1981), Moons (1982):
Analytical Theory of Lunar Rotation
- Herring (1986): VLBI Determination of
Nutation
- Williams (1980's): Numerical Theory of
Moon's Rotation, JPL/LE
- 1988~: Int'l Earth Rotation Service



1. Rigid Body

- Idealization of Finite Body
- ~ System of Discrete Mass points
 - Barycenter itself = Orbital Motion
 - Around Barycenter = Rotation
- Simplification $\mathbf{X}(t) = \mathbf{X}_o(t) + \mathbf{R}(t)\mathbf{x}$
 - Newtonian Mechanics
 - Constraint Conditions = Rigidity

Concept



- What is “rigid body”?
 - Set of points with **invariant** mutual distances
- Description of Position: Basic Triangle
 - Ex.: Center O, Two Points on A- and C-axis
 - Assuming Constraint Force
- Incompatible with Relativity (Why?)
- Idealization in Newton Mechanics (Why?)



Characteristics

- Def.: Body with Internal Invariant Distances
 - Or "System of Discrete Points under Constraints"
- How Many Freedom? (Answer = 6)
 - Ex.: 3 (Barycenter Position) + 2 (Direction of Rotation Axis) + 1 (Rotation Angle)
- Two Coordinate System (CS) are Needed
 - Inertial CS vs **Body-Fixed** CS
 - Ex.: Geocentric CS vs Earth-Crust-Fixed CS



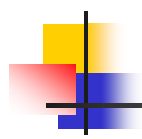
Number of Freedom = 6

- Outline of Proof
 - Assume that position of basic triangle is given
 - Number of freedom of arbitrary fourth point = 3
 - Distances between the fourth point and the basic triangle are known and fixed
 - 3 unknowns for 3 conditions = fourth point is fixed
 - Arbitrary fourth point is fixed = body is fixed
 - Number of freedom of basic triangle = $3 \times 3 = 9$
 - 3 distances between vertices are invariant
 - The rest number of freedom = $9 - 3 = 6$



Expression of Position

- Barycenter (3 positions)
- + Orientation (3 variables)
- Various Ways of Expression
 - Triad of Basis Vectors $\mathbf{E}=(\mathbf{e}_A \ \mathbf{e}_B \ \mathbf{e}_C)$
 - 3 Angles: Eulerian Angles (ψ, θ, ϕ)
 - Rotation Matrix $\mathbf{R}=\mathbf{R}(\psi, \theta, \phi)$
 - Quaternion $q_0 + q_i i + q_j j + q_k k$
 - Spinor $s_0 \sigma_0 + s_1 \sigma_1 + s_2 \sigma_2 + s_3 \sigma_3$



Expression of Motion

- Barycenter Orbital Motion: Linear Velocity
- Rigid Body Rotation: Angular Velocity
- Motion of Barycenter $\mathbf{X}_0(t)$
- Rigid Body Rotation
 - Finite Rotation = Rotation Matrix $\mathbf{R}(t)$
 - Infinitesimal Rotation = Vector Product

$$d\mathbf{e}_J = \mathbf{e}_J(t + dt) - \mathbf{e}_J(t) = \boldsymbol{\omega}(t) dt \times \mathbf{e}_J(t)$$



Equation of Motion

- Barycenter: Orbital Motion $\mathbf{p} = m\mathbf{v}$
 - Position \mathbf{x} , Velocity \mathbf{v} , Mass m , Momentum \mathbf{p}
 - Newton: Var. of Mom. = Force \mathbf{F} $\frac{d\mathbf{p}}{dt} = \mathbf{F}$
- Orientation: Rotation
 - Angle θ , Ang. Velocity ω , Moment of Inertia \mathbf{I} , Angular Momentum \mathbf{L} $\mathbf{L} = \mathbf{I}\omega$
 - Euler: Var. of Ang. Mom. = Torque \mathbf{N}
 - **Difference**: \mathbf{I} is a Matrix $\frac{d\mathbf{L}}{dt} = \mathbf{N}$



Solution

- Free (=No Torque) Rotation
 - Ang. Vel. (A=B): Simple, Trigonometric F.
 - Ang. Vel. (General): Complicated, Elliptic F.
 - Angle: Incomplete Elliptic Integrals
- Under Torque
 - Generally Torque is **Small** = Perturbation
 - Analytical Theory vs Numerical Integration
 - Euler Angles vs Canonical Variables

2. Mathematics of Rotation



- Basics of Kinematics of Rigid Body
 - Infinitesimal Rotation: Vector Product
 - Finite Rotation: Rotation Matrix
- Euler's Theorem
- Fundamental Rotation Matrix
- Angular Velocity Vector

Body-Fixed CS

- Two Distinct Coordinate System (CS)
 - Inertial CS (XYZ) vs Body-Fixed CS (ABC)
 - Ex. of Latter CS = Rigid Body CS
- Expression of CS $\mathbf{E} = (\mathbf{e}_A \ \mathbf{e}_B \ \mathbf{e}_C)$
 - Basis Matrix = **Triad** of Basis Vectors
 - Orthonormality Conditions (6) $\mathbf{e}_j \cdot \mathbf{e}_k = \delta_{jk}$
- Meaning of Vector Components

$$\mathbf{L} = L_X \mathbf{e}_X + L_Y \mathbf{e}_Y + L_Z \mathbf{e}_Z = L_A \mathbf{e}_A + L_B \mathbf{e}_B + L_C \mathbf{e}_C$$

Basis Vectors

- Completeness => Time Var.

$$\frac{d\mathbf{e}_j}{dt} = \sum_{k=1}^3 \Omega_{jk} \mathbf{e}_k$$

- Orthonormality => Anti-symmetry

$$\frac{d(\mathbf{e}_j \cdot \mathbf{e}_k)}{dt} = 0 \rightarrow \frac{d\mathbf{e}_j}{dt} \cdot \mathbf{e}_k + \mathbf{e}_j \cdot \frac{d\mathbf{e}_k}{dt} = 0 \rightarrow \Omega_{jk} = -\Omega_{kj}$$

- Multiplication of Anti-symmetric Matrix

- = **Vector Product**

$$\frac{d\mathbf{e}_A}{dt} = \boldsymbol{\omega} \times \mathbf{e}_A, \dots$$



Angular Velocity

- What is Vector ω ?
- Simple Ex.: Const. C-axis Comp. $\omega = \omega \mathbf{e}_C$
 - C-axis is invariant $\mathbf{e}_C(t) = \mathbf{e}_{C0}$
 - Other Axis rotates with constant speed
$$\mathbf{e}_A(t) = \mathbf{e}_{A0} \cos \omega t + \mathbf{e}_{B0} \sin \omega t$$
$$\mathbf{e}_B(t) = -\mathbf{e}_{A0} \sin \omega t + \mathbf{e}_{B0} \cos \omega t$$
- Vector $\omega =$ **Angular Velocity Vector**

Addition Theorem

- Add Two Rotations $\frac{d\mathbf{e}_A}{dt} = \boldsymbol{\omega}_1 \times \mathbf{e}_A, \dots$ $\frac{d\mathbf{e}_A}{dt} = \boldsymbol{\omega}_2 \times \mathbf{e}_A, \dots$

- Infinitesimal Rotations: δt is small

$$\mathbf{e}_A^{(1)} \approx \mathbf{e}_A^{(0)} + \boldsymbol{\omega}_1 \delta t \times \mathbf{e}_A^{(0)}$$

$$\mathbf{e}_A^{(2)} \approx \mathbf{e}_A^{(1)} + \boldsymbol{\omega}_2 \delta t \times \mathbf{e}_A^{(1)} \approx \mathbf{e}_A^{(0)} + (\boldsymbol{\omega}_1 + \boldsymbol{\omega}_2) \delta t \times \mathbf{e}_A^{(0)}$$

- Sum of Angular Velocity = **Vector Sum**

Finite Rotation

- Expressions

- Matrix, Spinor, Quaternions

- Operation = Matrix Product

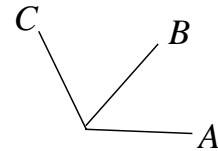
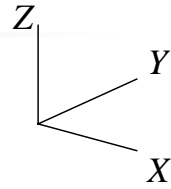
- Then, **order-dependent**


$E(t)$

- Change Basis $M.$ to Another Basis $M.$

- Rotation Matrix = **Orthogonal**

- 9 Comp. but Only 3 Independent





Rigid = Orthogonal

- Rigidity = Distance Invariant

$$\mathbf{X}^2 = \mathbf{x}^2$$

- Assuming Linear Transf. $\mathbf{X} = \mathbf{R}\mathbf{x}$

- Rigidity = Orthogonality

$$(\mathbf{R}\mathbf{x})^2 = \mathbf{x}^T \mathbf{R}^T \mathbf{R} \mathbf{x} = \mathbf{x}^2$$

$$\therefore \mathbf{R}^T \mathbf{R} = \mathbf{1} \quad \text{or} \quad \mathbf{R}^{-1} = \mathbf{R}^T$$

Angular Velocity M.

- Time Differentiation of Orthogonal M.
 - Orthogonality = Angular Velocity M.

$$\mathbf{R}^T \mathbf{R} = \mathbf{1} \rightarrow \mathbf{R}^T \frac{d\mathbf{R}}{dt} = - \left(\frac{d\mathbf{R}^T}{dt} \right) \mathbf{R} \equiv -\boldsymbol{\Omega}$$

- Anti-symmetry

$$\boldsymbol{\Omega}^T = -\boldsymbol{\Omega}$$

$$\boldsymbol{\Omega} = \begin{pmatrix} 0 & -\omega_z & \omega_y \\ \omega_z & 0 & -\omega_x \\ -\omega_y & \omega_x & 0 \end{pmatrix}$$

Vector Product

- Ang. Vel. M. = 3x3 Anti-Symmetric M.
- Equivalent with 3-Dim. **Axial** vector

$$\mathbf{\Omega} = \begin{pmatrix} 0 & -\omega_z & \omega_y \\ \omega_z & 0 & -\omega_x \\ -\omega_y & \omega_x & 0 \end{pmatrix} \quad \mathbf{\omega} = \begin{pmatrix} \omega_x \\ \omega_y \\ \omega_z \end{pmatrix}$$

- Anti-symmetric M. = Vector Product

$$\mathbf{\Omega}\mathbf{x} = \mathbf{\omega} \times \mathbf{x}$$



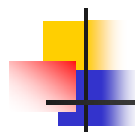
Euler's Theorem

- Arbitrary Finite Rotation = Screw Rotation
= **Triple Product** of Fund. Rotational M.

$$\mathbf{R} = \mathbf{R}_{ijk}(\alpha, \beta, \gamma) \equiv \mathbf{R}_k(\gamma)\mathbf{R}_j(\beta)\mathbf{R}_i(\alpha)$$

$$\left(\mathbf{R}_{ijk}(\alpha, \beta, \gamma)\right)^{-1} = \mathbf{R}_{kji}(-\gamma, -\beta, -\alpha)$$

- **Euler Angles** = 3 Fund. Rotation Angles
 - Also called as Cardano's angles

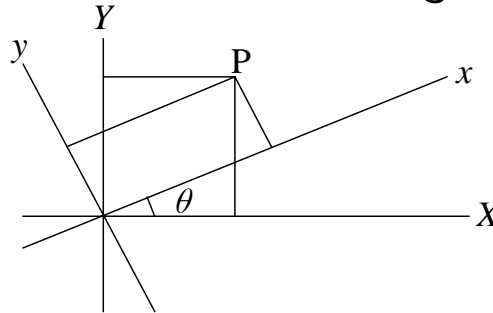


Fundamental Rotation

- Ex.: Third Axis Rotation with Angle θ

$$\mathbf{R}_3(\theta)$$

$$= \mathbf{R}_Z(\theta)$$



- Coord. Transf.

$$x_P = X_P \cos \theta + Y_P \sin \theta, \quad y_P = -X_P \sin \theta + Y_P \cos \theta$$

$$X_P = x_P \cos \theta - y_P \sin \theta, \quad Y_P = x_P \sin \theta + y_P \cos \theta$$

Properties of Fund. R.

- Rotation around j-th axis $\mathbf{R}_j(\theta)$
- Inverse = Transpose = Reverse Rotation

$$(\mathbf{R}_j(\theta))^{-1} = (\mathbf{R}_j(\theta))^T = \mathbf{R}_j(-\theta)$$

- Old Coord. \mathbf{X} , New Coord. \mathbf{x}

$$\begin{pmatrix} x_P \\ y_P \\ z_P \end{pmatrix} = \mathbf{R}_j(\theta) \begin{pmatrix} X_P \\ Y_P \\ Z_P \end{pmatrix}$$

$$\begin{pmatrix} X_P \\ Y_P \\ Z_P \end{pmatrix} = \mathbf{R}_j(-\theta) \begin{pmatrix} x_P \\ y_P \\ z_P \end{pmatrix}$$



Expression of Fund. R.

$$\mathbf{R}_1(\theta) = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta & \sin \theta \\ 0 & -\sin \theta & \cos \theta \end{pmatrix}$$

$$\mathbf{R}_2(\theta) = \begin{pmatrix} \cos \theta & 0 & -\sin \theta \\ 0 & 1 & 0 \\ \sin \theta & 0 & \cos \theta \end{pmatrix}$$

$$\mathbf{R}_3(\theta) = \begin{pmatrix} \cos \theta & \sin \theta & 0 \\ -\sin \theta & \cos \theta & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Two Different Rotations

- Two Different Expr. Of Same Vector

$$\mathbf{L} = L_X \mathbf{e}_X + L_Y \mathbf{e}_Y + L_Z \mathbf{e}_Z = L_A \mathbf{e}_A + L_B \mathbf{e}_B + L_C \mathbf{e}_C$$

- CS Rotation vs Rotation of Coordinates

$$\mathbf{E} = (\mathbf{e}_A \quad \mathbf{e}_B \quad \mathbf{e}_C) \quad \begin{pmatrix} L_A \\ L_B \\ L_C \end{pmatrix} = \mathbf{R} \begin{pmatrix} L_X \\ L_Y \\ L_Z \end{pmatrix}$$

- What is **Relation** between **E** and **R**?



Small Angle Approx.

- Assume All Angles are Small

$$\mathbf{R}_3(\theta) \cong \mathbf{1} + \begin{pmatrix} 0 & \theta & 0 \\ -\theta & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix} = \mathbf{1} + \begin{pmatrix} 0 \\ 0 \\ -\theta \end{pmatrix} \times = \mathbf{1} - \theta \mathbf{e}_3 \times$$

$$\therefore \prod_j \mathbf{R}_j(\theta_j) \cong \mathbf{1} - \left(\sum_j \theta_j \mathbf{e}_j \right) \times$$

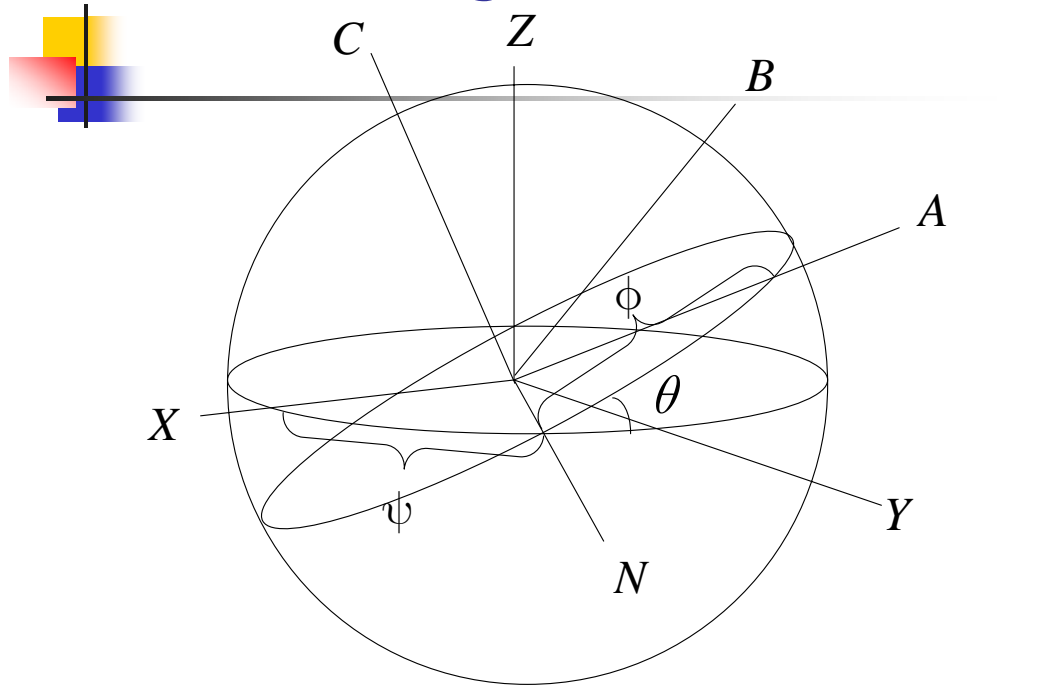


Euler Rotation

- Number of Combinations = $3 \times 2 \times 2$
 - Why 12 ways?
- 3-1-3 Sequence (= X convention)
 - Classic Dynamics, Continental Style
 - American Style changes ψ and ϕ

$$\mathbf{R}_{313}(\psi, \theta, \phi) = \mathbf{R}_3(\phi) \mathbf{R}_1(\theta) \mathbf{R}_3(\psi)$$

Euler Angles (3-1-3)





3-1-3 Rotation Matrix

- Expression in Inertial CS (Q: Derive)

$$\mathbf{R}_{313}(\psi, \theta, \phi) = \begin{pmatrix} C_\phi C_\psi - S_\phi C_\theta S_\psi & C_\phi S_\psi + S_\phi C_\theta C_\psi & S_\phi S_\theta \\ -S_\phi C_\psi - C_\phi C_\theta S_\psi & -S_\phi S_\psi + C_\phi C_\theta C_\psi & C_\phi S_\theta \\ S_\theta S_\psi & -S_\theta C_\psi & C_\theta \end{pmatrix}$$
$$= \begin{pmatrix} \cos \phi \cos \psi - \sin \phi \cos \theta \sin \psi & \cos \phi \sin \psi + \sin \phi \cos \theta \cos \psi & \sin \phi \sin \theta \\ -\sin \phi \cos \psi - \cos \phi \cos \theta \sin \psi & -\sin \phi \sin \psi + \cos \phi \cos \theta \cos \psi & \cos \phi \sin \theta \\ \sin \theta \sin \psi & -\sin \theta \cos \psi & \cos \theta \end{pmatrix}$$



3-1-3 Rotation Basis

- Transformation of Basis Vectors

$$(\mathbf{e}_X, \mathbf{e}_Y, \mathbf{e}_Z) \rightarrow (\mathbf{e}_N, \mathbf{e}_M, \mathbf{e}_Z) \rightarrow (\mathbf{e}_N, \mathbf{e}_P, \mathbf{e}_C) \rightarrow (\mathbf{e}_A, \mathbf{e}_B, \mathbf{e}_C)$$

- Expression of Basis Vectors (Q: Derive)

$$\mathbf{e}_N = \mathbf{e}_X \cos \psi + \mathbf{e}_Y \sin \psi, \quad \mathbf{e}_M = -\mathbf{e}_X \sin \psi + \mathbf{e}_Y \cos \psi$$

$$\mathbf{e}_P = \mathbf{e}_M \cos \theta + \mathbf{e}_Z \sin \theta, \quad \mathbf{e}_C = -\mathbf{e}_M \sin \theta + \mathbf{e}_Z \cos \theta$$

$$\mathbf{e}_A = \mathbf{e}_N \cos \phi + \mathbf{e}_P \sin \phi, \quad \mathbf{e}_B = -\mathbf{e}_N \sin \phi + \mathbf{e}_P \cos \phi$$



3-1-3 Rotation Basis (2)

- Relation of Basis and Rotation Matrices

$$\mathbf{E}^T = (\mathbf{e}_A \quad \mathbf{e}_B \quad \mathbf{e}_C)^T = \mathbf{R}_3(\phi)\mathbf{R}_1(\theta)\mathbf{R}_3(\psi)$$

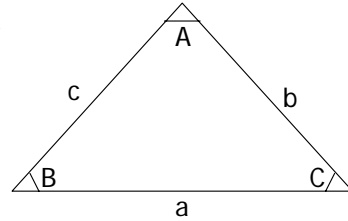
- Inertial CS Expr. of Basis (Q: Derive)

$$\mathbf{e}_N = \begin{pmatrix} \cos \psi \\ \sin \psi \\ 0 \end{pmatrix}, \mathbf{e}_M = \begin{pmatrix} -\sin \psi \\ \cos \psi \\ 0 \end{pmatrix}, \mathbf{e}_P = \begin{pmatrix} -\cos \theta \sin \psi \\ \cos \theta \cos \psi \\ \sin \theta \end{pmatrix}, \mathbf{e}_C = \begin{pmatrix} \sin \theta \sin \psi \\ -\sin \theta \cos \psi \\ \cos \theta \end{pmatrix}$$
$$\mathbf{e}_A = \begin{pmatrix} \cos \phi \cos \psi - \sin \phi \cos \theta \sin \psi \\ \cos \phi \sin \psi + \sin \phi \cos \theta \cos \psi \\ \sin \phi \sin \theta \end{pmatrix}, \mathbf{e}_B = \begin{pmatrix} -\sin \phi \cos \psi - \cos \phi \cos \theta \sin \psi \\ -\sin \phi \sin \psi + \cos \phi \cos \theta \cos \psi \\ \cos \phi \sin \theta \end{pmatrix}$$

Spherical Triangles

- Angles: A, B, C + Arcs: a, b, c
- 2 Expr. Of Same Rotation

$$\mathbf{R}_1(A)\mathbf{R}_3(c)\mathbf{R}_1(B) = \mathbf{R}_3(b)\mathbf{R}_1(\pi - C)\mathbf{R}_3(a)$$



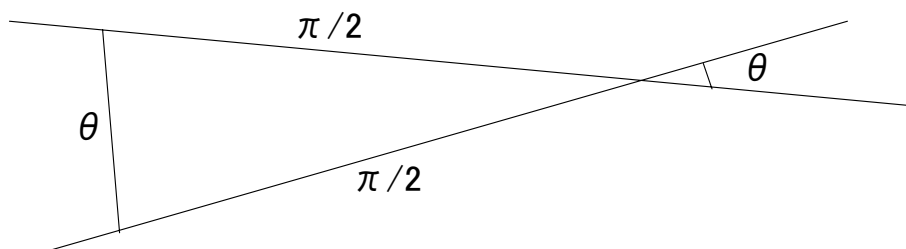
- **Formulas** of Spherical Triangles


- Sine $\sin a \sin B = \sin b \sin A$
- 1st Cosine $\cos a = \cos b \cos c + \sin b \sin c \cos A$
- 2nd Cosine $\cos A = -\cos B \cos C + \sin B \sin C \cos a$

Defect of 3-1-3 Seq.

- **Difficulty** in Expressing Axis-2 Rotation

$$\mathbf{R}_2(\theta) = \mathbf{R}_3\left(\frac{-\pi}{2}\right)\mathbf{R}_1(\theta)\mathbf{R}_3\left(\frac{\pi}{2}\right)$$



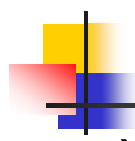


Defects of 3-1-3 Seq. (2)

- **Degeneracy** in case of small angles

$$\mathbf{R}_{313}(\psi, \theta, \phi) \cong \mathbf{1} - \begin{pmatrix} \theta \\ 0 \\ \phi + \psi \end{pmatrix} \times$$

- Cause: 2 indices are the same
- Solution
 - All-different-indices Seq. like 1-2-3 Seq.



3-2-3 Sequence

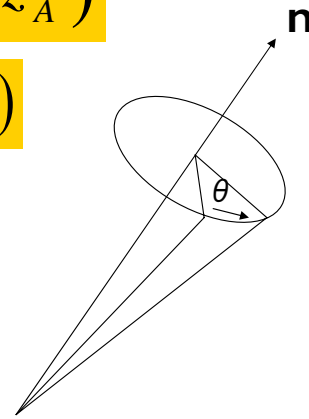
- = Y Convention = English Style

$$\mathbf{P} = \mathbf{R}_{323}(-\zeta_A, \theta_A, -z_A)$$

- Screw Rotation $\mathbf{R}_{323}(\lambda, \varphi, \theta)$

= Rotation around Axis \mathbf{n}

$$\mathbf{n} = \begin{pmatrix} \sin \varphi \cos \lambda \\ \sin \varphi \sin \lambda \\ \cos \varphi \end{pmatrix}$$





Screw Rotation

- Screw Rotation

- = Rodriguez's Rotation Matrix

$$\mathbf{X} = \mathbf{R}_{323}(\lambda, \varphi, \theta) \mathbf{x}$$

$$= \mathbf{x} + \sin \theta \mathbf{n} \times \mathbf{x} + (1 - \cos \theta) \mathbf{n} \times (\mathbf{n} \times \mathbf{x})$$

- Angular Velocity = Constant

$$\boldsymbol{\omega} = \omega \mathbf{n}$$

- Then, Rotation Angle is

$$\theta = \omega t$$



Other Sequences

- 1-3-1 Seq.: used in Nutation Matrix

$$\mathbf{N} = \mathbf{R}_{131} \left(\varepsilon_A, -\Delta\psi, -(\varepsilon_A + \Delta\varepsilon) \right)$$

- 2-1-3 Seq.: Polar Motion + Sidereal R.

$$\mathbf{W S} = \mathbf{R}_{312} \left(\Theta, -y_p, -x_p \right)$$



1-2-3 Sequence

- Aerodynamics, Attitude Dynamics
 - Mathematically **Best**

$$\mathbf{R}_{123}(\alpha, \beta, \gamma) = \mathbf{R}_3(\gamma)\mathbf{R}_2(\beta)\mathbf{R}_1(\alpha)$$

$$= \begin{pmatrix} C_\gamma C_\beta & C_\gamma S_\beta S_\alpha + S_\gamma C_\alpha & -C_\gamma S_\beta C_\alpha + S_\gamma S_\alpha \\ -S_\gamma C_\beta & -S_\gamma S_\beta S_\alpha + C_\gamma C_\alpha & S_\gamma S_\beta C_\alpha + C_\gamma S_\alpha \\ S_\beta & -C_\beta S_\alpha & C_\beta C_\alpha \end{pmatrix} \cong \mathbf{I} - \begin{pmatrix} \alpha \\ \beta \\ \gamma \end{pmatrix} \times$$



Angular Velocity Expr.

- Vector Expression of Infinitesimal Rotation

$$\boldsymbol{\omega} dt = d\boldsymbol{\theta} = \sum_j \mathbf{e}_{\theta_j} d\theta_j \rightarrow \boldsymbol{\omega} = \sum_j \left(\frac{d\theta_j}{dt} \right) \mathbf{e}_{\theta_j}$$

- Rotation Axis = Direction of Angular Velocity
 - Note 1: Ang. Velocity \neq Time Var. of Angles
 - Note 2: Basis Vector Change with Rotation
 - Note 3: Angular Velocity is independent on the order of Rotational Operations

An Example of Expr.

■ 3-1-3 Seq. $\boldsymbol{\omega} = \frac{d\psi}{dt} \mathbf{e}_\psi + \frac{d\theta}{dt} \mathbf{e}_\theta + \frac{d\phi}{dt} \mathbf{e}_\phi$

■ Inertial
CS Repr. $\mathbf{e}_\psi = \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix}, \mathbf{e}_\theta = \begin{pmatrix} \cos \psi \\ \sin \psi \\ 0 \end{pmatrix}, \mathbf{e}_\phi = \begin{pmatrix} \sin \theta \sin \psi \\ -\sin \theta \cos \psi \\ \cos \theta \end{pmatrix}$

■ Components $\begin{pmatrix} \omega_x \\ \omega_y \\ \omega_z \end{pmatrix} = \begin{pmatrix} \dot{\theta} \cos \psi + \dot{\phi} \sin \theta \sin \psi \\ \dot{\theta} \sin \psi - \dot{\phi} \sin \theta \cos \psi \\ \dot{\psi} + \dot{\phi} \cos \theta \end{pmatrix}$



Another Example

- Body-Fixed CS Repr.

$$\mathbf{e}_\psi = \begin{pmatrix} \sin \theta \sin \phi \\ \sin \theta \cos \phi \\ \cos \theta \end{pmatrix}, \mathbf{e}_\theta = \begin{pmatrix} \cos \phi \\ -\sin \phi \\ 0 \end{pmatrix}, \mathbf{e}_\phi = \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix}$$

- Components
- = Euler's Kinematic Eq. of Rotation

$$\begin{pmatrix} \omega_A \\ \omega_B \\ \omega_C \end{pmatrix} = \begin{pmatrix} \dot{\psi} \sin \theta \sin \phi + \dot{\theta} \cos \phi \\ \dot{\psi} \sin \theta \cos \phi - \dot{\theta} \sin \phi \\ \dot{\psi} \cos \theta + \dot{\phi} \end{pmatrix}$$

Velocity Transf.

- Variation of Basis Vector
- Transf. Law in Body-Fixed CS
- D/Dt = Diff. in Body-Fixed CS
- Proof

$$\frac{d\mathbf{e}_A}{dt} = \boldsymbol{\omega} \times \mathbf{e}_A, \dots$$

$$\mathbf{v} = \frac{D\mathbf{x}}{Dt} + \boldsymbol{\omega} \times \mathbf{x}$$

$$\mathbf{X} = x_A \mathbf{e}_A + x_B \mathbf{e}_B + x_C \mathbf{e}_C, \mathbf{V} = v_A \mathbf{e}_A + v_B \mathbf{e}_B + v_C \mathbf{e}_C$$

$$\mathbf{x} = (x_A, x_B, x_C)^T, \mathbf{v} = (v_A, v_B, v_C)^T$$

$$\mathbf{V} = \frac{d\mathbf{X}}{dt} = \frac{dx_A}{dt} \mathbf{e}_A + \frac{dx_B}{dt} \mathbf{e}_B + \frac{dx_C}{dt} \mathbf{e}_C + x_A \frac{d\mathbf{e}_A}{dt} + x_B \frac{d\mathbf{e}_B}{dt} + x_C \frac{d\mathbf{e}_C}{dt}$$

$$\therefore v_A = \frac{dx_A}{dt} + (\omega_B x_C - \omega_C x_B) \rightarrow \mathbf{v} = \frac{D\mathbf{x}}{Dt} + \boldsymbol{\omega} \times \mathbf{x}$$

Another Expr. of Transf.

- Symbolic Form $\frac{d}{dt} = \frac{D}{Dt} + \boldsymbol{\omega} \times$
- Two Time Diff.
 - True Time Diff. in Inertial CS: d/dt
 - Component in Body-Fixed CS : D/Dt
- Inertial CS vs Body-Fixed CS

$$\mathbf{V} = \frac{d\mathbf{X}}{dt}$$

$$\mathbf{v} = \frac{D\mathbf{x}}{Dt} + \boldsymbol{\omega} \times \mathbf{x}$$



3. Physics of Rotation

- Total Momentum & Total Angular Momentum
- Conservation of Angular Momentum
- Ang. Mom. Expr. in Body-Fixed CS
- Moment of Inertia
- Expression of Torque
- Euler's Dynamical Eq. of Rotation



Newton's Eq. of Motion

- For Each Component of Finite Body
- Total Force = External + **Internal**
- Internal Force: Molecular, Atomic, EM,...

$$m_K \frac{d^2 \mathbf{X}_K}{dt^2} = \mathbf{F}_K + \sum_{J \neq K} \mathbf{F}_{JK}$$

Motion of Barycenter

- Total Mass $M \equiv \sum_K m_K$
- Barycenter $M\mathbf{X}_O \equiv \sum_K m_K \mathbf{X}_K$
- Internal F.: Weak Law of Reaction $\mathbf{F}_{JK} = -\mathbf{F}_{KJ}$
- Barycenter Mot. independent on **Internal F.**

$$M \frac{d^2 \mathbf{X}_O}{dt^2} = \mathbf{F} \equiv \sum_K \mathbf{F}_K$$

Angular Momentum

■ Linear Mom. vs Ang. Mom. $\mathbf{L}_K \equiv m_K \mathbf{X}_K \times \mathbf{V}_K$

■ Total Ang. Mom.

= Orbital AM + Rotational AM

$$\mathbf{L}_{\text{Total}} \equiv \sum_K \mathbf{L}_K$$

■ Q: Derive

$$\mathbf{L}_{\text{Total}} = \mathbf{L}_O + \mathbf{L}, \quad \mathbf{L}_O = M \mathbf{X}_O \times \mathbf{V}_O,$$

$$\mathbf{L} = \sum_K m_K (\mathbf{X}_K - \mathbf{X}_O) \times (\mathbf{V}_K - \mathbf{V}_O)$$



Discrete to Continuous

- Discrete Sum \rightarrow Continuous Integration
- Mass Density Distribution Function $\rho(\mathbf{x})$

$$\sum_K m_K \mathbf{P}_K \rightarrow \int \mathbf{P}(\mathbf{x}) \rho(\mathbf{x}) d^3 \mathbf{x}$$

- Ex.: Rotational Angular Momentum, \mathbf{L}

$$\mathbf{L} = \sum_K m_K (\mathbf{X}_K - \mathbf{X}_O) \times (\mathbf{V}_K - \mathbf{V}_O)$$
$$\rightarrow \mathbf{L} = \int \mathbf{x} \times \mathbf{v}(\mathbf{x}) \rho(\mathbf{x}) d^3 \mathbf{x}$$

Total Angular Mom.

- Strong Law of Reaction
 - = Constraint F. is Central (Invalid for EM force)

$$\mathbf{F}_{JK} \propto \mathbf{X}_J - \mathbf{X}_K$$

- Time Var. of Total AM by External Force
 - Q: Show

$$\frac{d\mathbf{L}_{\text{Total}}}{dt} = \sum_K \mathbf{X}_K \times \mathbf{F}_K$$

Rotational Angular Mom.

- Accelerations $\mathbf{A}_O \equiv \frac{\mathbf{F}}{M}$ $\mathbf{A}_K \equiv \frac{\mathbf{F}_K}{m_K}$
- Barycenter O
- Time Var. of Rot. Ang. Mom. by **Torque**

$$\begin{aligned}\frac{d\mathbf{L}}{dt} &= \sum_K m_K (\mathbf{X}_K - \mathbf{X}_O) \times (\mathbf{A}_K - \mathbf{A}_O) \\ &= \int \mathbf{x} \times \mathbf{a}(\mathbf{x}) \rho(\mathbf{x}) d^3\mathbf{x} \equiv \mathbf{N}\end{aligned}$$

- Tidal Acceleration $\mathbf{a}_K = \mathbf{A}_K - \mathbf{A}_O$

Tidal Acceleration

- Def: **Difference** from Barycenter's Accel., \mathbf{A}_O
- Usual Approximation
 - Depend on Position Only
 - Up to 1st Order of Variation
- Ex.: Gravitational F. by External Mass point

$$\mathbf{a}_K \approx \left(\frac{\partial \mathbf{A}_K}{\partial \mathbf{x}_K} \right)_O \mathbf{x}_K$$

$$\mathbf{A}_K = \frac{Gm}{|\mathbf{r} - \mathbf{x}_K|^3} (\mathbf{r} - \mathbf{x}_K) \quad \mathbf{a}_K = \left(\frac{Gm}{r^5} \right) [3(\mathbf{r} \cdot \mathbf{x}_K) \mathbf{r} - r^2 \mathbf{x}_K]$$

- Note: Not always $\mathbf{A}_O = \mathbf{A}_K(\mathbf{x}_K = \mathbf{x}_O)$

Rotational Ang. Mom. (2)

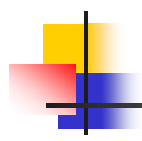
- Rigid Body Approx.: $D\mathbf{x}/Dt=0 \rightarrow \mathbf{v} = \boldsymbol{\omega} \times \mathbf{x}$
 - Rot. AM, \mathbf{L} , is Linear Function of $\boldsymbol{\omega}$

$$\mathbf{L} = \int \mathbf{x} \times (\boldsymbol{\omega} \times \mathbf{x}) \rho(\mathbf{x}) d^3\mathbf{x} = \mathbf{I}\boldsymbol{\omega}$$

- **Moment-of-Inertia Tensor \mathbf{I}** (Q: Show)

$$(\mathbf{I})_{ij} = I_{ij} = \int (\mathbf{x}^2 \delta_{ij} - x_i x_j) \rho(\mathbf{x}) d^3\mathbf{x}$$

- “AM Axis” = Direction of AM Vector



Moment of Inertia

- Mass Moment Tensor

$$J_{ij} = \int x_i x_j \rho(\mathbf{x}) d^3 \mathbf{x}$$

- Moment-of-Inertia Tensor (Q: Show)

$$I_{XX} = \frac{1}{2}(J_{YY} + J_{ZZ} - J_{XX}) \quad I_{ij} = -J_{ij} \quad (i \neq j)$$

- Cyclic Change of Indices (X→Y→Z)

Rotational Energy

- Again, Rigid Body Approx.: $\mathbf{v} = \boldsymbol{\omega} \times \mathbf{x}$

$$T = \frac{1}{2} \int \mathbf{v}^2 \rho(\mathbf{x}) d^3 \mathbf{x} = \frac{1}{2} \int (\boldsymbol{\omega} \times \mathbf{x})^2 \rho(\mathbf{x}) d^3 \mathbf{x}$$
$$= \frac{\boldsymbol{\omega} \cdot \mathbf{L}}{2} = \frac{\boldsymbol{\omega} \cdot (\mathbf{I} \boldsymbol{\omega})}{2} = \frac{1}{2} \sum_{i,j=1}^3 I_{ij} \omega_i \omega_j$$

- Maximum Energy for Constant AM
→ "Rotation Axis" = "AM Axis"

Ellipsoid of Inertia

- Normal Ellipsoid assoc. with $\sum_{i,j=1}^3 I_{ij}x_i x_j = 1$
- Quadratic Form: Diagonalization
 $\tilde{I}_{11}\tilde{x}_1^2 + \tilde{I}_{22}\tilde{x}_2^2 + \tilde{I}_{33}\tilde{x}_3^2 = 1$
- Diag. Elem. = **Principal Mom. Of Inertia**
 $A \leq B \leq C$ $Ax^2 + By^2 + Cz^2 = 1$
- Inverse $a \equiv A^{-1} \geq b \equiv B^{-1} \geq c \equiv C^{-1}$

Principal CS

- = Principal Moment-of-Inertia CS
- CS where \mathbf{I} is Diagonalized
- Well-used as a Body-Fixed CS
- Expr. in Principal CS
- **Figure Axis**
= C-Axis of Principal CS

$$\mathbf{I} = \begin{pmatrix} A & 0 & 0 \\ 0 & B & 0 \\ 0 & 0 & C \end{pmatrix}$$



Expr. In Principal CS

- Rotational Quant. Are **Plainly** Expressed

- Rotational AM

$$\begin{pmatrix} L_A \\ L_B \\ L_C \end{pmatrix} = \begin{pmatrix} A\omega_A \\ B\omega_B \\ C\omega_C \end{pmatrix}$$

- Rotational Energy

$$T = \frac{1}{2} (A\omega_A^2 + B\omega_B^2 + C\omega_C^2)$$



Triad of Principal CS

- In Inertial CS: 3-1-3 Euler Angles

$$(\mathbf{e}_A \quad \mathbf{e}_B \quad \mathbf{e}_C)^T = \mathbf{R}_3(\phi) \mathbf{R}_1(\theta) \mathbf{R}_3(\psi)$$

$$\mathbf{e}_A = \begin{pmatrix} \cos \psi \cos \phi - \sin \psi \cos \theta \sin \phi \\ \sin \psi \cos \phi + \cos \psi \cos \theta \sin \phi \\ \sin \theta \sin \phi \end{pmatrix}$$

$$\mathbf{e}_C = \begin{pmatrix} \sin \psi \sin \theta \\ -\cos \psi \sin \theta \\ \cos \theta \end{pmatrix}$$

$$\mathbf{e}_B = \begin{pmatrix} -\cos \psi \sin \phi - \sin \psi \cos \theta \cos \phi \\ -\sin \psi \sin \phi + \cos \psi \cos \theta \cos \phi \\ \sin \theta \cos \phi \end{pmatrix}$$

Moment of Inertia (2)

- **Dyadic** Expr. by Triad of Principal CS

$$\mathbf{I} = A\mathbf{e}_A \otimes \mathbf{e}_A + B\mathbf{e}_B \otimes \mathbf{e}_B + C\mathbf{e}_C \otimes \mathbf{e}_C$$

- Easy Inverse Matrix Repr. (Q: Show)

$$\mathbf{I}^{-1} = a\mathbf{e}_A \otimes \mathbf{e}_A + b\mathbf{e}_B \otimes \mathbf{e}_B + c\mathbf{e}_C \otimes \mathbf{e}_C$$

- Dyadic = Direct Product of Vector

$$(\mathbf{a} \otimes \mathbf{b})_{jk} = (\mathbf{a})_j (\mathbf{b})_k = a_j b_k$$



Principal Mom. Inertia

- Spherically Symmetric
 - $A=B=C$
- Rot. Symm. (Oblate or Hamburger-like)
 - $A=B<C$
- Rot. Symm. (Prolate or Lemon-shaped)
 - $A<B=C$
- General: $A<B<C$

Principal M. Inertia (2)

- Uniform Triaxial Ellipsoid (Q: Confirm)
 - Cyclic Change of Indices (A→B→C)

$$A = \frac{4\pi}{15} R_A R_B R_C (R_B^2 + R_C^2) \quad \frac{C - A}{B} = \frac{R_A^2 - R_C^2}{R_A^2 + R_C^2}$$

- Ex: Asteroid Ida, A<B~C

$$R_A = 59.8 \text{ km}, R_B = 25.4 \text{ km}, R_C = 18.6 \text{ km}$$

$$\frac{A}{C} = 0.235, \frac{B}{C} = 0.929$$

Principal M. Inertia (3)

- Earth: $A \sim B < C$

$$\frac{C_E - A_E}{C_E} = 3.284741 \times 10^{-3}, \frac{B_E - A_E}{C_E} = 2.196 \times 10^{-6}, \frac{C_E}{M_E R_E^2} = 0.330701$$

$$GM_E = 3.986005 \times 10^{14} \text{ m}^3 \text{ s}^{-2}, R_E = 6.3781366 \times 10^6 \text{ m}$$

- Moon: $A < B < C$

$$\frac{C_M - A_M}{B_M} = 6.32 \times 10^{-4}, \frac{B_M - A_M}{C_M} = 2.33 \times 10^{-4}, \frac{C_M}{M_M R_M^2} = 0.389$$

$$\frac{M_M}{M_E} = 1.230002 \times 10^{-2}, R_M = 1.738 \times 10^6 \text{ m}$$

Three Axes

- Rotat. $\boldsymbol{\omega}$ vs Ang.Mom. \mathbf{L} vs Figure \mathbf{e}_C
 - Confusing
 - Which is so-called "Pole"?
- Expressions in Principal CS

$$\boldsymbol{\omega} = \begin{pmatrix} \omega_A \\ \omega_B \\ \omega_C \end{pmatrix}, \mathbf{L} = \begin{pmatrix} A\omega_A \\ B\omega_B \\ C\omega_C \end{pmatrix}, \mathbf{e}_C = \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix}$$

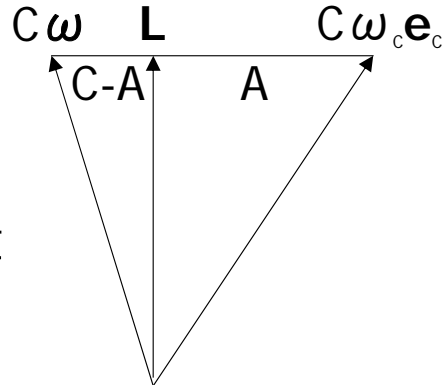
Three Axes (2)

- If $A=B < C$ $\mathbf{L} = A\boldsymbol{\omega} + (C - A)\omega_c \mathbf{e}_c$

- Q: Show

- 3 Axes are on the Same Plane

- If $A \sim C$, AM Axis is close to Rot. Axis but far from Figure Axis





Visualization of Rotation

- Method of Poinsot (1851)
 - Free Rotation: Moving and Fixed Cones
- Polhode
 - Orbit of Rot. Axis on Ellipsoid of Inertia
 $\omega_A(t), \omega_B(t), \omega_C(t)$ $A\omega_A^2 + B\omega_B^2 + C\omega_C^2 = 2T$
- Herpolhode
 - Orbit of Rot. Axis on Plane in Inertial CS
 $\omega_X(t), \omega_Y(t), \omega_Z(t)$ $c_X\omega_X + c_Y\omega_Y + c_Z\omega_Z = c_0$



Eq. of Rotation

- Conservation of Rotational Ang. Mom.

- Inertial CS + Ang. Mom. $\frac{d\mathbf{L}}{dt} = \mathbf{N}$

- Principal CS + Ang. Mom.

- $\frac{D\mathbf{L}}{Dt} = \mathbf{N} - (\mathbf{I}^{-1}\mathbf{L}) \times \mathbf{L}$

- Principal CS + Ang. Vel. (Euler 1765)

- $\frac{D\boldsymbol{\omega}}{Dt} = \mathbf{I}^{-1} (\mathbf{N} - \boldsymbol{\omega} \times \mathbf{I}\boldsymbol{\omega})$



Tidal Torque

- **Main Term** of External Torque

$$\mathbf{N} = \frac{3\mu}{r^5} \int (\mathbf{r} \cdot \mathbf{x})(\mathbf{r} \times \mathbf{x}) \rho(\mathbf{x}) d^3 \mathbf{x}$$

- Ex.: Moon+Sun Acting on Earth
 - Moon (or Sun)'s GM = μ , **Geocentric** Position \mathbf{r}
- General Torque: from Law of Reaction
 - F = Force by Body on External Point

$$\mathbf{N} = -\mathbf{r} \times \mathbf{F}$$

Eq. of Rotation (2)

■ Inertial CS $(\psi, \theta, \phi, \mathbf{L})$

- Kinematic Part = Expr. of Ang. Vel. with Time Var. of Euler Angles

$$\frac{d\psi}{dt} = \omega_z - \left(\frac{d\phi}{dt}\right) \cos \theta$$

$$\frac{d\mathbf{L}}{dt} = \mathbf{N}$$

$$\frac{d\theta}{dt} = \omega_x \cos \psi + \omega_y \sin \psi$$

$$\boldsymbol{\omega} = \mathbf{I}^{-1}\mathbf{L}$$

$$\frac{d\phi}{dt} = \frac{\omega_x \sin \psi - \omega_y \cos \psi}{\sin \theta}$$

$$\mathbf{N} = N_A \mathbf{e}_A + N_B \mathbf{e}_B + N_C \mathbf{e}_C$$

Eq. of Rotation (3)

■ Principal CS, Ang. Mom. $(\psi, \theta, \phi, L_A, L_B, L_C)$

■ Q: Show

$$\frac{d\psi}{dt} = \frac{aL_A \sin \phi + bL_B \cos \phi}{\sin \theta}$$

$$\frac{d\theta}{dt} = aL_A \cos \phi - bL_B \sin \phi$$

$$\frac{d\phi}{dt} = cL_C - \left(\frac{d\psi}{dt} \right) \cos \theta$$

$$\frac{dL_A}{dt} = N_A - (b - c) L_B L_C$$

$$\frac{dL_B}{dt} = N_B - (c - a) L_C L_A$$

$$\frac{dL_C}{dt} = N_C - (a - b) L_A L_B$$

Eq. of Rotation (4)

■ Principal CS, Ang. Vel. $(\psi, \theta, \phi, \omega_A, \omega_B, \omega_C)$

■ Q: Show

$$\frac{d\psi}{dt} = \frac{\omega_A \sin \phi + \omega_B \cos \phi}{\sin \theta}$$

$$\frac{d\theta}{dt} = \omega_A \cos \phi - \omega_B \sin \phi$$

$$\frac{d\phi}{dt} = \omega_C - \left(\frac{d\psi}{dt} \right) \cos \theta$$

$$\frac{d\omega_A}{dt} = \frac{N_A - (C - B)\omega_B\omega_C}{A}$$

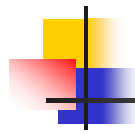
$$\frac{d\omega_B}{dt} = \frac{N_B - (A - C)\omega_C\omega_A}{B}$$

$$\frac{d\omega_C}{dt} = \frac{N_C - (B - A)\omega_A\omega_B}{C}$$



Trivial Solution: $A=B=C$

- Simplest Solution: Spherically Symmetric
- Zero Torque (Q: Why?)
- \rightarrow Const. Ang. Mom. \rightarrow Const. Ang. Vel.
- $A=B=C \rightarrow$ AM Axis = Rot. Axis
- Coord. Transf.: New C-Axis = AM Axis
- Time Variation of Angles
 - Rot. Angle is Linear, Others = Const.
 - Result: **Uniform Rotation**



Orbit vs Rotation

Item	Orbit	Rotation
Position	X	Euler A., ($\psi \theta \phi$)
Velocity	V	Ang. Vel., ω
Mass	M	Mom. Inertia, I
Momentum	P=MV	L=I ω
Kin. Energy	$MV^2/2$	$\omega \cdot I \omega / 2$
Eq. of Motion	$d\mathbf{P}/dt=\mathbf{F}$	$d\mathbf{L}/dt=\mathbf{N}$
Free Motion	Linear Motion	Free Rotation



4. Free Rotation

- = **Torque-free** (vs Keplerian Motion)
 - Separation of Ang. Vel. and Angles → Integrable
- Integrals: Energy, T ; Ang. Mom., (L_x, L_y, L_z)
- Special Solution
 - Rotation Around one of Principal Axes
- General Solution
 - $A=B$: Trig. F., Free Polar Motion & Precession
 - $A < B$: Elliptic F. + Incomplete Elliptic Integrals



Special Solution

- AM Axis = one of Principal Axes
- Solution = Uniform Rotation
- Case of C-axis
 - Dynamics Part $L_A = L_B = 0, L_C = C\omega_0$
 - Kinematics Part $\psi = \psi_0, \theta = \theta_0, \phi = \phi_0 + \omega_0 t$
 - After Some Coord. Transf. $\psi = \theta = 0, \phi = \omega_0 t$
- Stability?

Stability of Special Sol.

- Assumption: Other Ang. Vel. Comp. are Small
- Linear Stability Theory: Growth Rate Eval.

$$L_A, L_B \ll L_C \sim L_0 \quad L_A, L_B \propto e^{\lambda t}$$

$$\frac{dL_A}{dt} \sim -(b-c)L_0L_B, \quad \frac{dL_B}{dt} \sim -(c-a)L_0L_A$$

$$\rightarrow \lambda^2 = (b-c)(c-a)L_0^2 < 0$$

- Cyclic Change of Indices (A,a \rightarrow B,b \rightarrow C,c)
- Stable: A- or C-axis \Leftrightarrow Unstable: B-axis

A=B<C

- Non-trivial Integrals

- C-axis Component of **L**

$$\frac{dL_C}{dt} = 0 \rightarrow L_C = C\omega_0$$


- A- and B-Comp. of **L**: **Harmonic Oscillation**

$$\frac{dL_A}{dt} = -\Omega L_B$$

$$\Omega = (a - c) L_C = \frac{C - A}{A} \omega_0$$

$$\frac{dL_B}{dt} = +\Omega L_A$$

$$\left(\frac{2\pi}{\Omega} \right)_{\text{Earth}} \approx 303.438 \text{ day}$$



A=B<C (2)

- View in Body-Fixed CS: \mathbf{L} = Slow Precession

$$L_A = L_{A0} \cos \Omega t - L_{B0} \sin \Omega t$$

$$L_B = L_{A0} \sin \Omega t + L_{B0} \cos \Omega t$$

- Similar Harm Osc. For Ang. Vel. Ω

= Free Polar Motion

$$\omega_A = \omega_{A0} \cos \Omega t - \omega_{B0} \sin \Omega t$$

= Euler Motion

$$\omega_B = \omega_{A0} \sin \Omega t + \omega_{B0} \cos \Omega t$$

- Earth's Euler Motion: $P \sim 303$ days

A=B<C (3)

- New Z-axis = AM Axis

$$L_A = G \sin \theta \sin \phi$$

$$L_B = G \sin \theta \cos \phi$$

$$L_C = G \cos \theta$$

$$G \equiv |\mathbf{L}|$$

$$\phi = \tan^{-1} \left(\frac{L_A}{L_B} \right) = \phi_0 - \Omega t, \quad \theta = \theta_0 \equiv \cos^{-1} \left(\frac{L_C}{G} \right)$$

- Rot. Angle ϕ , Inclination θ -> Precession A. ψ

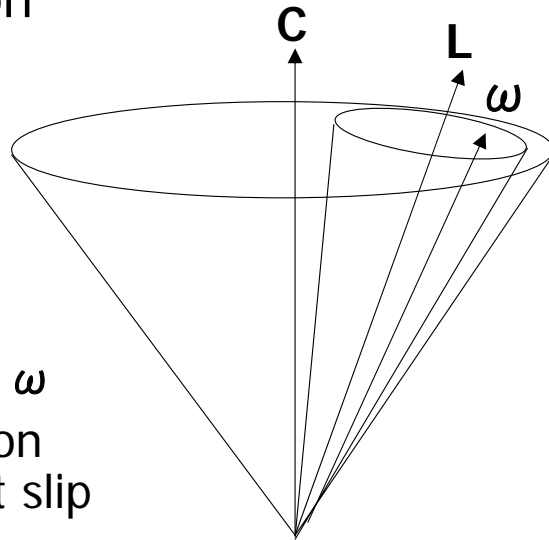
$$\frac{d\psi}{dt} = \frac{1}{\cos \theta} \left(\omega_0 - \frac{d\phi}{dt} \right) \rightarrow \psi = \psi_0 + aGt$$

- Solution: **Slow Reverse Rotation + Constant Inclination + Fast Precession**



Moving and Fixed Cones

- Poincot's Expression
= Geometric View in
Body-Fixed CS
($A=B < C$)
 - Fixed C.: C-axis
 - Moving C.: L-axis
 - Modulus: Rot. Axis ω
 - Moving Cone rolls on
Fixed Cone without slip





$A < B < C$

- Eq. of Free Rotation (Principal CS + AM)

- Two **Conservatives**
(Q: Confirm)

$$2T = aL_A^2 + bL_B^2 + cL_C^2$$
$$G^2 = L_A^2 + L_B^2 + L_C^2$$

$$\frac{dL_A}{dt} = -(b - c)L_B L_C$$

$$\frac{dL_B}{dt} = -(c - a)L_C L_A$$

$$\frac{dL_C}{dt} = -(a - b)L_A L_B$$

$A < B < C$ (2)

- Principal CS + Ang. Vel.

- Parameters of Inertia

$$\alpha \equiv \frac{C-B}{A} > 0, \beta \equiv \frac{A-C}{B} < 0, \gamma \equiv \frac{B-A}{C} > 0$$

- Identity on Param. Of Inertia (Q: Show)

$$\alpha + \beta + \gamma + \alpha\beta\gamma = 0$$



$A < B < C$ (3)

- Eq. of Rotation (Principal CS + Ang. Vel.)

- Two Conservatives

Again (Q: Show)

$$2T = A\omega_A^2 + B\omega_B^2 + C\omega_C^2$$

$$G^2 = A^2\omega_A^2 + B^2\omega_B^2 + C^2\omega_C^2$$

$$\frac{d\omega_A}{dt} = -\alpha\omega_B\omega_C$$

$$\frac{d\omega_B}{dt} = -\beta\omega_C\omega_A$$

$$\frac{d\omega_C}{dt} = -\gamma\omega_A\omega_B$$

$A < B < C$ (4)

- Characteristics (Depend on Initial Cond.)
 - Charact. Mom. Inertia $D \equiv \frac{G^2}{2T}$ $A \leq D \leq C$
 - Inverse Charact. Mol $d \equiv \frac{1}{D}$ $a \geq d \geq c$
= Inverse of Charact. Mol
 - Charact. Ang. Vel. $\omega_D \equiv \frac{2T}{G}$
 - Relation $D\omega_D = G$

$A < B < C$ (5)

- Kirchhoff Transformation
- Using Characteristics

- Time: $t \rightarrow u$ $u \equiv G \sqrt{(a-d)(b-c)}(t-t_0)$

- Ang. Mom.: $L_A, L_B, L_C \rightarrow x, y, z$

$$\frac{L_A}{G} \equiv \sqrt{\frac{d-c}{a-c}}x, \quad \frac{L_B}{G} \equiv \sqrt{\frac{d-c}{b-c}}y, \quad \frac{L_C}{G} \equiv \sqrt{\frac{a-d}{a-c}}z$$

- Initial Condition $L_B = 0$ when $t = t_0$

$A < B < C$ (6)

- Transformed Eq. of Motion (Q: Confirm)
 - Q: Derive Kirchhoff Transf. s.t. Resulting This Form

$$\frac{dx}{du} = -yz, \quad \frac{dy}{du} = zx, \quad \frac{dz}{du} = -k^2 xy$$

- **Modulus** of Elliptic F.: k

$$k^2 \equiv \frac{(a-b)(d-c)}{(a-d)(b-c)} \geq 0$$

- Jacobi's Elliptic Func.

$$x = \operatorname{cn}(u; k), \quad y = \operatorname{sn}(u; k), \quad z = \operatorname{dn}(u; k)$$

$A < B < C$ (7)

- **Low Energy** Assumption ($b < d \rightarrow k < 1$)

- New Z-axis
= **L** axis

$$L_A = A\omega_A = G \sin \theta \sin \phi = G \sqrt{\frac{d-c}{a-c}} \operatorname{cn}(u; k)$$

$$L_B = B\omega_B = G \sin \theta \cos \phi = G \sqrt{\frac{d-c}{b-c}} \operatorname{sn}(u; k)$$

- Polhode

$$L_C = C\omega_C = G \cos \theta = G \sqrt{\frac{a-d}{a-c}} \operatorname{dn}(u; k)$$

- AM Ellipse

$$(a-c)L_A^2 + (b-c)L_B^2 = (d-c)G^2$$

- AV Ellipse

$$(C-A)A\omega_A^2 + (C-B)B\omega_B^2 = 2CT - G^2$$

$A < B < C$ (8)

- Solution of Rotation Angle and Inclination
 - Wobbling and Slowly-decreasing ϕ (Q: Show)

$$\phi = \tan^{-1} \left(\frac{L_A}{L_B} \right) = \frac{\pi}{2} - \text{am}(u; k) - \tan^{-1} \left(\frac{(1-\sigma) \text{sn}(u; k) \text{cn}(u; k)}{\sigma + (1-\sigma) \text{sn}^2(u; k)} \right)$$

$$\sigma \equiv \sqrt{\frac{b-c}{a-c}} \leq 1$$

- Oscillating Inclination θ (Q: Show)

$$\theta = \cos^{-1} \left(\frac{L_C}{G} \right) = \cos^{-1} \left(\sqrt{\frac{a-d}{a-c}} \text{dn}(u; k) \right)$$

$A < B < C$ (9)

- Precession Angle ψ : Incompl. EI of 3rd Kind
 - Kinoshita (1992, CMDA)

$$\psi = \int \frac{aL_A \sin \phi + bL_B \cos \phi}{\sin \theta} dt$$
$$= \psi_0 + cG(t - t_0) + \frac{a - c}{\sqrt{(a - d)(b - c)}} \text{pn}(u; n, k)$$

- **Parameter** of Elliptic Integral

$$n \equiv \frac{a - b}{b - c} = \frac{1 - \sigma^2}{\sigma^2}$$

A < B < C (10)

- Period of Rotation Angle

$$P_\phi \equiv \frac{4K}{\nu} = \frac{4K}{G\sqrt{(a-d)(b-c)}}$$

$$\beta \equiv \tan^{-1} \sqrt{\frac{d-c}{a-d}}$$

- Period of Precession Angle (Q: Show)

$$P_\psi = \frac{2\pi}{G \left[c + (a-c) \left\{ \frac{k^2}{k^2+n} + \left(\frac{\pi}{2K} \right) \frac{\sqrt{n}\Lambda_0(\beta;k)}{\sqrt{(1+n)(k^2+n)}} \right\} \right]}$$

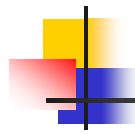
5. Elliptic Functions

- Jacobi's Elliptic Functions (Jacobi 1829): sn, cn, dn
 - "Bible": Byrd and Friedman (1954)

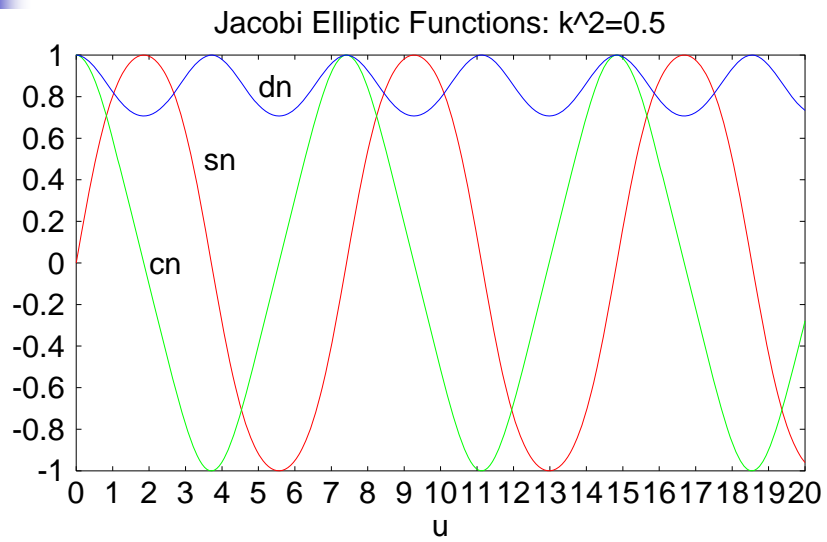
- sn: Inverse Function of Incomplete Elliptic Integral of 1st Kind in Rational Form $u = \int_0^x \frac{ds}{\sqrt{(1-s^2)(1-k^2s^2)}} \equiv \text{sn}^{-1}x$

$$\text{cn}(u; k) \equiv \pm \sqrt{1 - \text{sn}^2(u; k)}, \text{dn}(u; k) \equiv \sqrt{1 - k^2 \text{sn}^2(u; k)}$$

- Argument: u, Modulus: k
 - Usually We **Omit** Modulus



Graphs of Elliptic F.





Amplitude Function

- Inverse Function of Incomplete Elliptic Integral of 1st Kind in Trigonometric Form

$$u = \int_0^\varphi \frac{d\theta}{\sqrt{1 - k^2 \sin^2 \theta}} \equiv \operatorname{am}^{-1} \varphi$$

- **Amplitude** Function $\varphi = \operatorname{am}(u; k)$
- Trigonometric Expr. of Elliptic Functions

$$\operatorname{sn} u = \sin \varphi, \operatorname{cn} u = \cos \varphi, \operatorname{dn} u = \sqrt{1 - k^2 \sin^2 \varphi}$$



Modulus

- Modulus: k
- But, $m=k^2$ is more suitable
 - Clear from Definition of Incomplete EI

- **Complimentary** Modulus

$$k' = k_c \equiv \sqrt{1-k^2}$$

- Caution: Expression of k or m

$$\operatorname{sn}(u | m) \equiv \operatorname{sn}(u; k), \text{ etc.}$$



Formulas

- **Two Identities**

$$\operatorname{sn}^2 u + \operatorname{cn}^2 u = 1, \operatorname{dn}^2 u + k^2 \operatorname{sn}^2 u = 1$$

- **Special Modulus**

- $k=0$ $\operatorname{sn} u, \operatorname{cn} u, \operatorname{dn} u \rightarrow \sin u, \cos u, 1$

- $k=1$ $\operatorname{sn} u, \operatorname{cn} u, \operatorname{dn} u \rightarrow \tanh u, \operatorname{sech} u, \operatorname{sech} u$

- **Special Argument: $u=K/2$**

- $K(k)$ is Complete
EI of 1st Kind $(\operatorname{sn} u, \operatorname{cn} u, \operatorname{dn} u) = \left(\frac{1}{\sqrt{1+k}}, \sqrt{\frac{k}{1+k}}, \sqrt{k} \right)$



Notation

- Guderman Notation
 - Sine Amplitude=sn, Cosine Amplitude=cn
 - Delta Amplitude=dn, $dn(u;k) = \Delta(\phi)$
- Glaisher (**fraction**) notation
 - p,q,r = one of (s,c,d,n)
 - Examples

$$pq(u;k) \equiv \frac{pr(u;k)}{qr(u;k)}$$

$$nd(u;k) = \frac{1}{dn(u;k)}, sd(u;k) = \frac{sn(u;k)}{dn(u;k)}, cs(u;k) = \frac{cn(u;k)}{sn(u;k)}$$



Inverse Function

- Base: 2-Argument Inverse Elliptic Function

$$x = \operatorname{rcn}(u; k), y = \operatorname{rsn}(u; k) \rightarrow u = \operatorname{atn2}(y, x; k)$$

- Using Incomplete EI of 1st Kind in Trig. Form

$$\operatorname{atn2}(y, x; k) = F(\operatorname{atan2}(y, x); k)$$

- **General** Expression of Amplitude Function

$$\operatorname{am}(u; k) = 2\pi \left[\frac{u}{4K(k)} \right]_{\text{Gauss}} + \operatorname{atan2}(\operatorname{sn}(u; k), \operatorname{cn}(u; k))$$



Reciprocal Transf.

- Reduction of Modulus' Domain

- Reciprocal & Imaginary Transf. $\rightarrow 0 < k < 1$
- Gauss & Landen Transf. $\rightarrow k \sim 0$ or $k \sim 1$

- **Reciprocal**
Transformation

- $k > 1$ to $k < 1$

$$k \rightarrow k^{-1}$$

$$\operatorname{sn}(u; k) = k^{-1} \operatorname{sn}(ku; k^{-1})$$

$$\operatorname{cn}(u; k) = \operatorname{dn}(ku; k^{-1})$$

$$\operatorname{dn}(u; k) = \operatorname{cn}(ku; k^{-1})$$



Imaginary Transf.

- **Imaginary**

Transformation

- Pure Imaginary
k to Real k

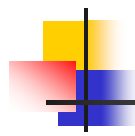
$$k \rightarrow \tilde{k} \equiv \sqrt{\frac{-k^2}{1-k^2}}$$

$$u \rightarrow \tilde{u} \equiv \sqrt{1-k^2}u$$

$$\operatorname{cn}(u; k) = \frac{\operatorname{cn}(\tilde{u}; \tilde{k})}{\operatorname{dn}(\tilde{u}; \tilde{k})}$$

$$\operatorname{sn}(u; k) = \frac{\sqrt{1-\tilde{k}^2} \operatorname{sn}(\tilde{u}; \tilde{k})}{\operatorname{dn}(\tilde{u}; \tilde{k})}$$

$$\operatorname{dn}(u; k) = \frac{1}{\operatorname{dn}(\tilde{u}; \tilde{k})}$$



Landen's Transf.

- (Ascending) **Landen's** Transformation

- Closer to 1

$$k \rightarrow \bar{k} \equiv \frac{2\sqrt{k}}{1+k}$$

$$u \rightarrow \bar{u} \equiv \frac{u}{1 + \sqrt{1 - \bar{k}^2}}$$

$$\text{cn}(u; k) = \frac{(1 + \sqrt{1 - \bar{k}^2}) \text{sn}(\bar{u}; \bar{k}) \text{cn}(\bar{u}; \bar{k})}{\text{dn}(\bar{u}; \bar{k})}$$

$$\text{sn}(u; k) = \frac{1 - (1 + \sqrt{1 - \bar{k}^2}) \text{sn}^2(\bar{u}; \bar{k})}{\text{dn}(\bar{u}; \bar{k})}$$

$$\text{dn}(u; k) = \frac{1 - (1 - \sqrt{1 - \bar{k}^2}) \text{sn}^2(\bar{u}; \bar{k})}{\text{dn}(\bar{u}; \bar{k})}$$



Gauss' Transformation

- = Descending Landen's Transf.
 - Closer to 0

$$k \rightarrow \hat{k} \equiv \left(\frac{k^2}{1 + \sqrt{1 - k^2}} \right)^2$$

$$u \rightarrow \hat{u} \equiv \frac{u}{1 + \hat{k}}$$

$$\text{cn}(u; k) = \frac{\text{cn}(\hat{u}; \hat{k}) \text{dn}(\hat{u}; \hat{k})}{1 + \hat{k} \text{sn}^2(\hat{u}; \hat{k})}$$

$$\text{sn}(u; k) = \frac{(1 + \hat{k}) \text{sn}(\hat{u}; \hat{k})}{1 + \hat{k} \text{sn}^2(\hat{u}; \hat{k})}$$

$$\text{dn}(u; k) = \frac{1 - \hat{k} \text{sn}^2(\hat{u}; \hat{k})}{1 + \hat{k} \text{sn}^2(\hat{u}; \hat{k})}$$

Jacobi's Nome, q

- Useful in Fast Computation
- Def.: K , and its Compliment, K' $K'(k) \equiv K(k')$

$$q \equiv \exp\left(\frac{-\pi K'}{K}\right) = \lambda \left(1 + 2\lambda^4 + 15\lambda^8 + 150\lambda^{12} + \dots\right)$$

- q is **small** when $k^2 < 1/2$ $k \leq \frac{1}{\sqrt{2}} \rightarrow q \leq e^{-\pi} \sim 0.0432$
- $k^2 > 1/2 \rightarrow$ Dual Transf.

- Expansion Factor $\lambda \equiv \frac{k^2}{2(1+k')(1+\sqrt{k'})^2} \approx \frac{k^2}{16}$ $k' \equiv \sqrt{1-k^2}$



Nome Expression

- Fraction Expr. Using Nome (**F**ast **C**omputation)
 - After Reduction to Basic Domain: $0 < k^2 < 1/2$

$$\operatorname{sn}(u; k) = \frac{\beta s y_s}{y_n}, \operatorname{cn}(u; k) = \frac{\sqrt{k'} \beta c y_c}{y_n}, \operatorname{dn}(u; k) = \frac{\sqrt{k'} y_d}{y_n}$$

$$s = \sin \nu, c = \cos \nu, \nu \equiv \frac{\pi u}{2K(k)} \quad \beta \equiv \frac{2\sqrt[4]{q}}{\sqrt{k}} = \frac{1 + 2(q + q^4 + q^9 + \dots)}{1 + q^2 + q^6 + q^{12} + \dots}$$

- Q: Draw Graphs of β & k as Function of q
- Normalized Elliptic Theta Functions: $y(q)$



q-Exp. of y-Functions

- Normalized **Theta Function**: Rapidly Converge

$$y_s = 1 - C_2 q^2 + C_4 q^6 - C_6 q^{12} + \dots$$

$$y_c = 1 + S_3 q^2 + S_5 q^6 + S_7 q^{12} + \dots$$

$$y_d = 1 + 2(T_2 q + T_4 q^4 + T_6 q^9 + \dots)$$

$$y_n = 1 - 2(T_2 q - T_4 q^4 + T_6 q^9 + \dots)$$

- Coefficients (= Chebychef Polynomials)

- Recursively Computable $C_n \equiv \frac{\sin(n+1)\nu}{\sin \nu}, S_n \equiv \frac{T_n}{\cos \nu}, T_n \equiv \cos(n\nu)$



Fourier Expansion

- Not-so-rapid convergence as q-expansion

$$\operatorname{sn}(u; k) = \frac{2\pi}{kK(k)} \sum_{n=0}^{\infty} \frac{q^{n+1/2}}{1-q^{2n+1}} \sin(2n+1)v$$

$$\operatorname{cn}(u; k) = \frac{2\pi}{kK(k)} \sum_{n=0}^{\infty} \frac{q^{n+1/2}}{1+q^{2n+1}} \cos(2n+1)v$$

$$\operatorname{dn}(u; k) = \frac{\pi}{2K(k)} + \frac{2\pi}{K(k)} \sum_{n=1}^{\infty} \frac{q^n}{1+q^{2n}} \cos 2nv$$

$$\operatorname{am}(u; k) = v + \sum_{n=1}^{\infty} \frac{2q^n}{n(1+q^{2n})} \sin 2nv$$

$$v \equiv \frac{\pi u}{2K(k)}$$



Numerical Computation

- Gauss' **Arithmetic-Geometric-Mean (AGM)**

- Common Conv. of Two Var.

$$a_{n+1} = \frac{a_n + b_n}{2}, \quad b_{n+1} = \sqrt{a_n b_n}$$

- Landen's Transf. of Ampl. F.

$$u_{n+1} = \frac{u_n + v_n}{2}, \quad v_{n+1} = \sqrt{u_n v_n}, \quad \phi_{n+1} = \frac{\phi_n + \sin^{-1}(k_n \sin \phi_n)}{2} \quad v_n \equiv k_n u_n$$

- Algorithm (Q: Write a Computer Code)

- Start ($u_0 = u, v_0 = ku$) and Repeat Transf. while Save $k_j = v_j / u_j$
 - After Convergence: $u_n = v_n$, i.e. $k_n = 1$, then $\phi_n = u_n$
 - Start (ϕ_n) and Repeat Reverse Transf. up to ϕ_0 using k_j
 - Details: Numerical Recipe, 2nd Ed.

Incomplete Elliptic Integrals

- Jacobi (1850): Indefinite Integral of **quartic rational**

- 1st Kind

$$F(x; k) = \int_0^x \frac{ds}{\sqrt{(1-s^2)(1-k^2s^2)}}$$

- 2nd Kind

$$E(x; k) = \int_0^x \sqrt{\frac{1-s^2}{1-k^2s^2}} ds$$

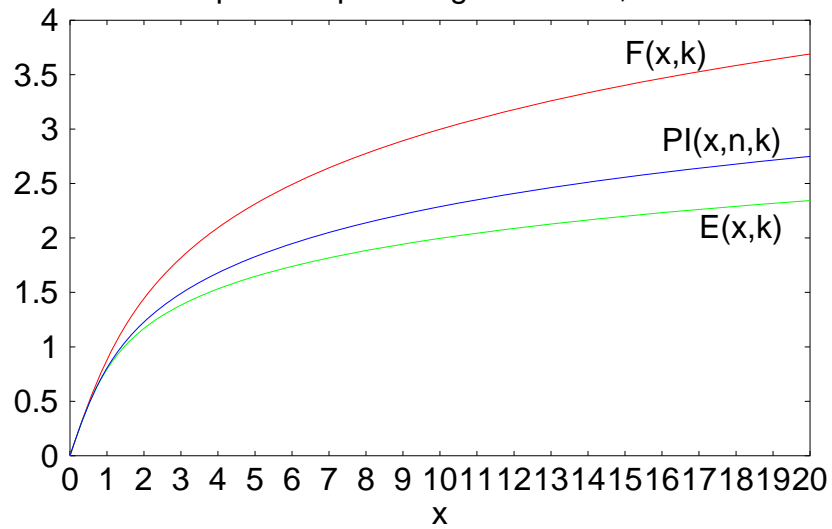
- 3rd Kind: Parameter, $n = -\alpha^2$

$$\Pi(x; n, k) \equiv \int_0^x \frac{ds}{(1+ns^2)\sqrt{(1-s^2)(1-k^2s^2)}}$$



Incomplete EI (2)

Incomplete Elliptic Integrals: $n=0.5$, $k^2=0.5$





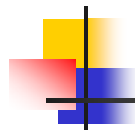
Incomplete EI (3)

- **Trigonometric** Representation (Legendre)
 - Most Popular, Note: Difference in Argument

- 1st Kind $F(\varphi; k) = \int_0^\varphi \frac{d\theta}{\sqrt{1-k^2 \sin^2 \theta}}$

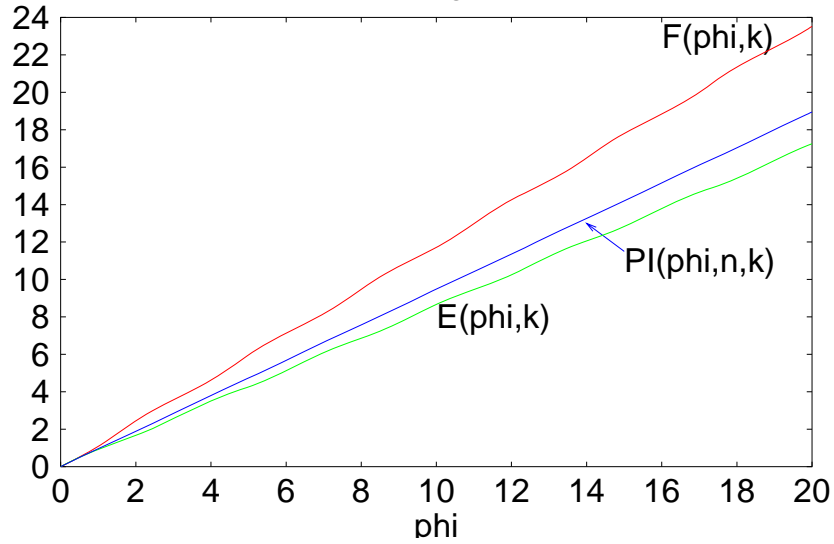
- 2nd Kind $E(\varphi; k) = \int_0^\varphi \frac{\cos^2 \theta d\theta}{\sqrt{1-k^2 \sin^2 \theta}}$

- 3rd Kind $\Pi(\varphi; n, k) = \int_0^\varphi \frac{d\theta}{(1+n \sin^2 \theta) \sqrt{1-k^2 \sin^2 \theta}}$



Incomplete EI (4)

Incomplete Elliptic Integrals: $n=0.5$, $k^2=0.5$





Incomplete EI (5)

- **Elliptic Function** Expression (Jacobi)

- Note: Difference in Argument
- Introducing New Symbols: en, pn

- 1st Kind $F(u; k) = u$

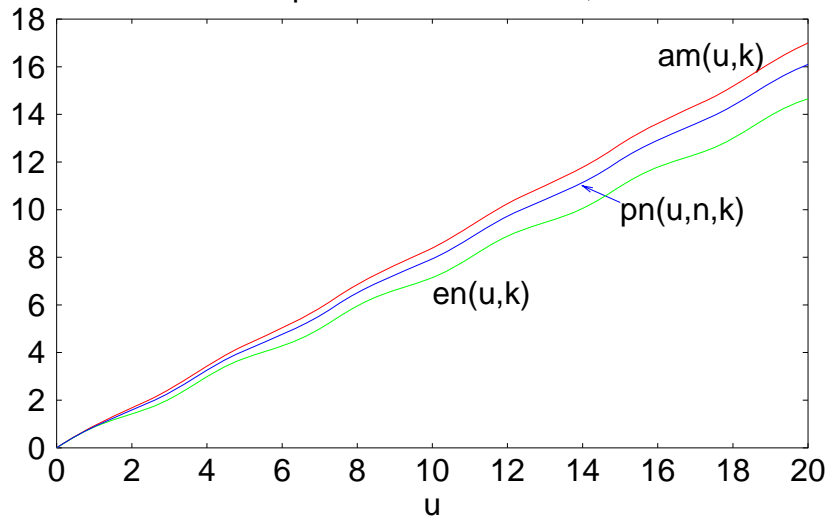
- 2nd Kind $E(u; k) = \int_0^u \text{cn}^2(v; k) dv = \text{en}(u; k)$

- 3rd Kind $\Pi(u; n, k) \equiv \int_0^u \frac{dv}{1 + n \text{sn}^2(v; k)} = \text{pn}(u; n, k)$



Incomplete EI (6)

Jacobi Elliptic Functions: $n=0.5$, $k^2=0.5$



Incomplete EI (7)

- Another Expr. Of Incomplete EI of 2nd Kind

$$E(u; k) = \left(\frac{K(k)}{E(k)} \right) u + Z(u; k)$$

- Jacobi's Zeta F. = Periodic Part of E(u;k)

$$Z(u; k) \equiv \text{zn}(u; k)$$

$$\approx \frac{\pi}{K(k)} \left(\frac{2q \sin 2v - 4q^4 \sin 4v + 9q^9 \sin 6v}{1 - 2q \cos 2v + 2q^4 \cos 4v - 2q^9 \cos 6v} \right) \quad v \equiv \frac{\pi u}{2K(k)}$$

Incomplete EI (8)

Special Case

- $k=0$

$$E(\varphi; 0) = F(\varphi; 0) = \varphi$$

$$\Pi(\varphi; n, 0) = \frac{\tan^{-1}(\sqrt{1+n} \tan \varphi)}{\sqrt{1+n}}$$

- $k=1$

$$E(\varphi; 1) = \sin \varphi \quad F(\varphi; 1) = \ln \left(\frac{1 + \sin \varphi}{\cos \varphi} \right)$$

- $n=-k^2$

$$\Pi(\varphi; n, 1) = \frac{1}{n+1} \left[n \tan^{-1}(n \sin \varphi) + \frac{1}{2} \log \left| \frac{1 - \sin \varphi}{1 + \sin \varphi} \right| \right]$$

$$\Pi(\varphi; -k^2, k) = \frac{1}{1-k^2} \left[E(\varphi; k) - \frac{k^2 \sin \varphi \cos \varphi}{\sqrt{1-k^2 \sin^2 \varphi}} \right]$$



Partial by Argument

- Partial Derivative w.r.t. Argument, u (Q: Show)
- Fixing Modulus, k , and Parameter, n

$$\frac{\partial \operatorname{sn}(u; k)}{\partial u} = \operatorname{cn} u \operatorname{dn} u$$

$$\frac{\partial \operatorname{cn}(u; k)}{\partial u} = -\operatorname{sn} u \operatorname{dn} u$$

$$\frac{\partial \operatorname{dn}(u; k)}{\partial u} = -k^2 \operatorname{sn} u \operatorname{cn} u$$

$$\frac{\partial \operatorname{am}(u; k)}{\partial u} = \operatorname{dn} u$$

$$\frac{\partial \operatorname{en}(u; k)}{\partial u} = \operatorname{cn}^2 u$$

$$\frac{\partial \operatorname{pn}(u; n, k)}{\partial u} = \frac{1}{1 + n \operatorname{sn}^2(u; k)}$$



Partial by Modulus

- Partial Derivative w.r.t. $m=k^2$ (Q: Show)

$$\frac{\partial \operatorname{sn}(u|m)}{\partial m} = \operatorname{cn}(u|m) \frac{\partial \operatorname{am}(u|m)}{\partial m}$$

$$\frac{\partial \operatorname{cn}(u|m)}{\partial m} = -\operatorname{sn}(u|m) \frac{\partial \operatorname{am}(u|m)}{\partial m}$$

$$\frac{\partial \operatorname{dn}(u|m)}{\partial m} = \frac{-\operatorname{sn}(u|m)}{2\operatorname{dn}(u|m)} \left[\operatorname{sn}(u|m) + 2\operatorname{cn}(u|m) \frac{\partial \operatorname{am}(u|m)}{\partial m} \right]$$

$$\frac{\partial \operatorname{am}(u|m)}{\partial m} = \operatorname{dn}(u|m) \left[\frac{\operatorname{pn}(u; -m|m) - u}{2m} \right]$$



Partial by Parameter

- Partial Derivatives w.r.t. m or n (Q: Show)

$$\frac{\partial \text{en}(u | m)}{\partial m} = \frac{\text{en}(u | m) - u}{2m}$$

$$\frac{\partial \text{pn}(u; n | m)}{\partial m} = \frac{1}{2} \left[u + \frac{\text{pn}(u; -m | m) - \text{pn}(u; n | m)}{m + n} \right]$$

$$\begin{aligned} \frac{\partial \text{pn}(u; n | m)}{\partial n} = & \frac{1}{2(1+n)} \left[\frac{\text{pn}(u; n | m) - u}{n} + \frac{1-m}{m+n} \text{pn}(u; -m | m) \right. \\ & \left. + \frac{\text{sn}(u | m) \text{cn}(u | m)}{\{1 + n \text{sn}^2(u | m)\} \text{dn}(u | m)} \right] \end{aligned}$$

Numerical Evaluation of Incomplete EI

- Numer. Evaluation of **General** Incomplete EI
(Fukushima and Ishizaki 1994b, CMDA)

$$G(\varphi; n_c, m_c, a, b) \equiv \int_0^{\varphi} \frac{a \cos^2 \theta + b \sin^2 \theta}{(\cos^2 \theta + n_c \sin^2 \theta) \sqrt{\cos^2 \theta + m_c \sin^2 \theta}} d\theta$$

- Basic Expression Formulas (Q: Confirm)

$$\lambda F(\varphi; k) + \mu E(\varphi; k) = G(\varphi; 1, 1 - k^2, \lambda + \mu, \lambda + \mu(1 - k^2))$$

$$\lambda F(\varphi; k) + \mu \Pi(\varphi; n, k) = G(\varphi; 1 + n, 1 - k^2, \lambda + \mu, \lambda(1 + n) + \mu)$$



General Elliptic F.

- Elliptic F. based on **General Incomplete EI**

$$\text{gn}(u; n | m, a, b) \equiv G(\text{am}(u | m); 1+n, 1-m, a, b)$$

- Basic Expression Formulas

$$\lambda u + \mu n \text{en}(u | m) = \text{gn}(u; 1, 1-m, \lambda + \mu, \lambda + \mu(1-m))$$

$$\lambda u + \mu p n \text{pn}(u; n | m) = \text{gn}(u; 1+n, 1-m, \lambda + \mu, \lambda(1+n) + \mu)$$

- Application No.1: (Q: Confirm)

$$\text{zn}(u | m) = \text{gn}\left(u; 1, 1-m, 1-\frac{K}{E}, 1-m-\frac{K}{E}\right)$$



Application of General EF

- Application No.2 (Q: Confirm)
 - Needed in Partial Computations

$$\frac{(1-m)u - en(u|m)}{m} = gn(u; 1, 1-m, 1, 0)$$

$$\frac{u - pn(u; n|m)}{n} = gn(u; 1+n, 1-m, 0, 1)$$

$$fn(u|m) \equiv \frac{pn(u; -m|m) - u}{m} = gn(u; 1-m, 1-m, 0, 1)$$



Complete EI

- Complete EI = Definite Integral

- 1st Kind $K(k) \equiv F(1;k) = \int_0^1 \frac{dx}{\sqrt{(1-x^2)(1-k^2x^2)}}$

- 2nd Kind $E(k) \equiv E(1;k) = \int_0^1 \sqrt{\frac{1-x^2}{1-k^2x^2}} dx$

- 3rd Kind

$$\Pi(n,k) \equiv \Pi(1;n,k) = \int_0^1 \frac{ds}{(1+ns^2)\sqrt{(1-s^2)(1-k^2s^2)}}$$

Complete EI (2)

- Another Expr. Of Complete EI of 3rd Kind ($n > 0$)

$$\Pi(n, k) = \frac{k^2 K(k)}{k^2 + n} + \left(\frac{\pi}{2}\right) \frac{\sqrt{n} \Lambda_0(\beta; k)}{\sqrt{(1+n)(k^2 + n)}} \quad \beta \equiv \tan^{-1} \frac{\sqrt{n}}{k}$$

- Heuman's Lambda Function

$$\Lambda_0(\beta; k) \equiv \frac{2}{\pi} \left[\{E(k) - K(k)\} F(\beta; k') + K(k) E(\beta; k') \right]$$



Complete EI (3)

- (Real) Period of Elliptic Functions: $4K$
 - $K(0) = \pi/2, K(1) = \infty$
- **Complimentary** Relations
 - $K'(k) \equiv K(k'), E'(k) \equiv E(k')$ $k' \equiv \sqrt{1-k^2}$
- Omitting Modulus k Frequently: K, K', E, E'
- Legendre's Relation

$$\Lambda_0\left(\frac{\pi}{2}, k\right) = 1 \rightarrow EK' + E'K - KK' = \frac{\pi}{2}$$



Dual Transformation

- = Complimentary Transf. = Prime Transf.

$$k \rightarrow k' \equiv \sqrt{1-k^2} \quad m \rightarrow m' \equiv 1-m$$

- $0 < k^2 < 1/2 \rightarrow 1/2 < k'^2 < 1$

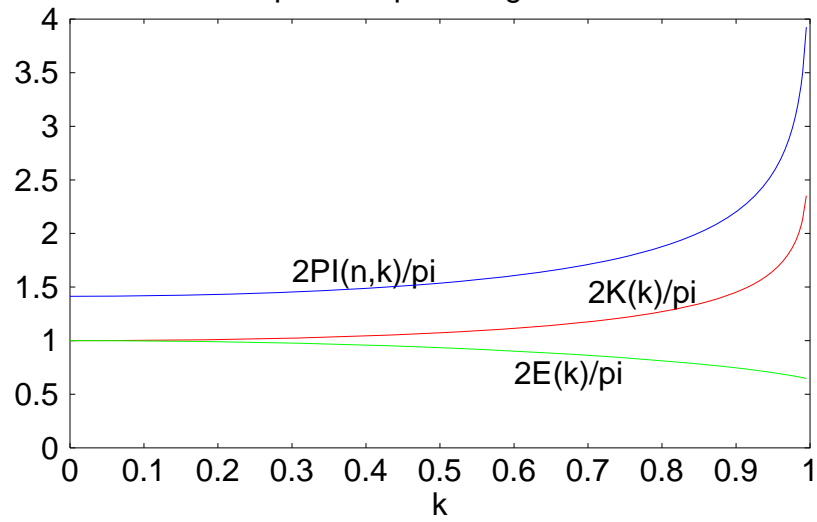
- Complete Elliptic Integrals $K, E \rightarrow K', E'$

- Nome $q \rightarrow q' \equiv \exp\left(\frac{\pi^2}{\log q}\right)$



Complete EI (2)

Complete Elliptic Integrals: $n=0.5$



Numerical Evaluation of Complete EI

- Inefficient: Calling Routines of Incomplete EI
- Standard: Using AGM (Numerical Recipe, etc)
- Practical: Modulus Transf. + q-Expansion
- Ex: $k^2 < 1/2$

- 1st Kind

$$K(k) = 2\pi \left(\frac{1 + 2q^4 + \dots}{1 + \sqrt{k'}} \right)^2$$

- 2nd Kind

$$E(k) = K(k) - \left(\frac{2\pi^2}{K(k)} \right) \left(\frac{q - 4q^4 + 9q^9 - \dots}{1 - 2(q - q^4 + q^9 - \dots)} \right)$$

Small Modulus

- Earth/Moon: $m \sim 0$ $k_{\text{Earth}} \sim 3 \times 10^{-8}, k_{\text{Moon}} \sim 5 \times 10^{-6}$
- Enough for 1st Order Exp. w.r.t. m

$$q \approx \frac{m}{16} \quad k' \approx 1 - \frac{m}{2} \quad K \approx \frac{\pi}{2} \left(1 + \frac{m}{2} \right) \quad E \approx \frac{\pi}{2} \quad v \approx \left(1 - \frac{m}{2} \right) u$$

$$\text{snu} = \left(1 + \frac{m}{4} \cos^2 u \right) \sin v \quad \text{cnu} = \left(1 - \frac{m}{4} \sin^2 u \right) \cos v$$

$$\text{dnu} = 1 - \frac{m}{2} \sin^2 u \quad \text{znu} \approx \frac{m}{4} \sin 2u \quad \text{enu} = \left(1 + \frac{m}{2} \right) u + \text{znu}$$



6. Forced Rotation

- Under Torque
 - Ex: Earth Precession/Nutation, Moon's Rotation
- Strength of Perturbation
 - Comparison of Rot. Energy vs Torque
- Unperturbed State
 - Simplification: $A=B < C$, $C-A \ll C$
 - Uniform Rotation around C-axis
- Type of Perturbation
 - Figure Axis Motion, Variation of Rot. Speed

Torque Expression

- In Inertial CS: **Complicated** (Q: Confirm)

$$N_X = \frac{3\mu}{r^5} [r_X r_Y I_{ZX} - r_Z r_X I_{XY} + (r_Y^2 - r_Z^2) I_{YZ} + r_Y r_Z (I_{ZZ} - I_{YY})]$$

- Cyclic Change of Indices (X→Y→Z)

- Moment-of-Inertia Components

$$I_{XY} = A(\mathbf{e}_A)_X (\mathbf{e}_A)_Y + B(\mathbf{e}_B)_X (\mathbf{e}_B)_Y + C(\mathbf{e}_C)_X (\mathbf{e}_C)_Y$$

$$\mathbf{I} = A\mathbf{e}_A \otimes \mathbf{e}_A + B\mathbf{e}_B \otimes \mathbf{e}_B + C\mathbf{e}_C \otimes \mathbf{e}_C$$

Torque Expression (2)

- In Principal CS: **Plain**

$$\begin{pmatrix} N_A \\ N_B \\ N_C \end{pmatrix} = \left(\frac{3\mu}{r^5} \right) \begin{pmatrix} (C-B)r_B r_C \\ (A-C)r_C r_A \\ (B-A)r_A r_B \end{pmatrix}$$

$$N_A = \mathbf{N} \cdot \mathbf{e}_{A,\dots}$$

$$r_A = \mathbf{r} \cdot \mathbf{e}_{A,\dots}$$

Torque Magnitude

- Dim. Of Torque = Energy
- Ratio of Rotational Energy and Torque
 - Earth: **Tiny**

$$\frac{N_E}{T_E} \approx \frac{3(C_E - A_E)}{C_E \omega_E^2} \left(\frac{\mu_S}{r_{SE}^3} + \frac{\mu_M}{r_{ME}^3} \right) \sin 2\varepsilon_E \approx 1.680 \times 10^{-7}$$

- Moon: Small $\varepsilon_M \approx 1^\circ 32' 33''$

$$\frac{N_M}{T_M} \approx \frac{3(C_M - A_M)}{C_M \omega_M^2} \left(\frac{\mu_E}{r_{EM}^3} \right) \sin 2\varepsilon_M \approx 9.91 \times 10^{-5}$$

A=B<C

- $N_C=0 \rightarrow$ Integral Exists
 - C-comp. of Ang. Mom.

$$\frac{dL_C}{dt} = 0 \rightarrow L_C = G_0$$

- Torque

$$\mathbf{N} = \frac{3\mu(C-A)r_C}{r^5} \mathbf{r} \times \mathbf{e}_C$$

- Poisson Approximation

$$L_A = L_B = 0 \rightarrow \mathbf{L} = L_C \mathbf{e}_C$$

- Eq. of Motion of ψ, θ

$$\frac{d\mathbf{e}_C}{dt} = \frac{\mathbf{N}}{G_0}$$

A=B<C (2)

- Eq. of Rotation in Inertial CS

$$\mathbf{e}_C = \begin{pmatrix} \sin \theta \sin \psi \\ -\sin \theta \cos \psi \\ \cos \theta \end{pmatrix}$$

$$G_0 \frac{d}{dt} \begin{pmatrix} \sin \theta \sin \psi \\ -\sin \theta \cos \psi \\ \cos \theta \end{pmatrix} = \begin{pmatrix} N_X \\ N_Y \\ N_Z \end{pmatrix}$$

- Eq. of Euler Angles in Poisson Approx.

- Z-comp. of Torque

Unnecessary (Why?)

$$\frac{d\phi}{dt} = \omega_C - \left(\frac{d\psi}{dt} \right) \cos \theta$$

$$\frac{d\psi}{dt} = \frac{N_X \cos \psi + N_Y \sin \psi}{G_0 \sin \theta}$$

$$\frac{d\theta}{dt} = \frac{N_X \sin \psi - N_Y \cos \psi}{G_0 \cos \theta}$$

$A \sim B < C$

- Laplacian Approximation

- Ignore $(A-B)/C$ in case of $A < B < C$

$$\mathbf{N} \simeq \frac{3\mu(2C - A - B)r_C}{2r^5} \mathbf{r} \times \mathbf{e}_C$$

- Dyn. Flattening, H + Poisson Approx.

$$\frac{C - A}{C} \rightarrow H \equiv \frac{2C - A - B}{2C}$$



Planar Circular Motion

- Pert. Body = **Circular Orbit on XY Plane**

- Ex: Sun's Pert. On Earth Rotation

- L: Helioc. Long. of Earth $r_c = \mathbf{r} \cdot \mathbf{e}_c = \bar{r} \sin \theta \sin(L - \psi)$

$$\mathbf{r} = -\bar{r} \begin{pmatrix} \cos L \\ \sin L \\ 0 \end{pmatrix}$$

$$\mathbf{r} \times \mathbf{e}_c = \bar{r} \begin{pmatrix} -\cos \theta \sin L \\ \cos \theta \cos L \\ \sin \theta \cos(L - \psi) \end{pmatrix}$$

- Torque in Inertial CS (Circular Approx.)

$$\begin{pmatrix} N_x \\ N_y \end{pmatrix} = \left[\frac{3\mu(C - A)}{2\bar{r}^3} \right] \sin 2\theta \sin(L - \psi) \begin{pmatrix} -\sin L \\ \cos L \end{pmatrix}$$

Precession Solution

- Eq. of Prec. Angle + Inclination (Circular)

$$\frac{d\psi}{dt} = -\kappa \cos \theta [1 - \cos 2(L - \psi)] \quad \kappa = \frac{3\mu(C - A)}{2C\omega_{c0}\bar{r}^3}$$

$$\frac{d\theta}{dt} = -\kappa \sin \theta \sin 2(L - \psi) \quad \psi^{(1)}(t) = \psi_0 - pt$$

- 1st Approx.: Precession $\theta^{(1)}(t) = \theta_0$
- Precession Const. $p = \kappa \cos \theta_0$
- Initial Cond. (Earth) $\psi_0 = 0, \theta_0 = -\varepsilon_0 < 0$



Luni-solar Precession

- Earth $\omega_{C_0} \approx 6.30039 \text{ rad/day}$ $\frac{C_E - A_E}{C_E} \approx 3.28474 \times 10^{-3}$
 $\cos \varepsilon_0 \approx 0.917482$
- Moon $\frac{\mu_M}{\bar{r}_{EM}^3} \approx \left(\frac{M_M}{M_E + M_M} \right) n_\ell^2 \approx 6.31786 \times 10^{-4} \text{ (rad/day)}^2$
- Sun $\frac{\mu_S}{\bar{r}_{SE}^3} \approx n_\ell^2 \approx 2.95908 \times 10^{-4} \text{ (rad/day)}^2$
- Luni-solar Prec.: **Crude** Estimate $\kappa \approx 37.2229''/\text{jy}$
 $p \approx (\kappa + \kappa') \cos \varepsilon_0 \approx 50.1467''/\text{jy}$ $\kappa' \approx 17.4340''/\text{jy}$



Short Period Nutation

- 2nd Approx. = **Main** Short Period Nutation

$$\psi^{(2)}(t) = \psi_0 - pt + \delta_1^{(2)}\psi \quad \theta^{(2)}(t) = \theta_0 + \delta_1^{(2)}\theta$$

$$\delta_1^{(2)}\psi = N_{2L} \cos \theta_0 \sin 2(L - \psi^{(1)})$$

$$\delta_1^{(2)}\theta = N_{2L} \sin \theta_0 \cos 2(L - \psi^{(1)})$$

$$N_{2L} = \frac{\kappa}{2(n_L + p)} \quad n_L = \frac{dL}{dt}$$

Elliptic Orbit

- Pert. Body: **Elliptic Orbit** in XY Plane
 - Mean Orb. Long. L Mean Anomaly ℓ
 - Pericenter Long. $\varpi \equiv L - \ell$
 - Semi-major Axis a Eccentricity e
- **Ignoring** 2nd and Higher Order of e

$$r_x = a \left(\cos L + \frac{e}{2} \cos(L + \ell) - \frac{3e}{2} \cos \varpi \right) \quad r_z = 0$$

$$r_y = a \left(\sin L + \frac{e}{2} \sin(L + \ell) - \frac{3e}{2} \sin \varpi \right) \quad r = a(1 - e \cos \ell)$$



Elliptic Orbit (2)

$$r_c = \mathbf{r} \cdot \mathbf{e}_c = -a \sin \theta \left[\sin(L - \psi) + \frac{e}{2} \sin(L - \psi + \ell) - \frac{3e}{2} \sin(\varpi - \psi) \right]$$

$$\mathbf{r} \times \mathbf{e}_c = a \begin{pmatrix} \cos \theta \left[\sin L + \frac{e}{2} \sin(L + \ell) - \frac{3e}{2} \sin \varpi \right] \\ -\cos \theta \left[\cos L + \frac{e}{2} \cos(L + \ell) - \frac{3e}{2} \cos \varpi \right] \\ -\sin \theta \left[\cos(L - \psi) + \frac{e}{2} \cos(L - \psi + \ell) - \frac{3e}{2} \cos(\varpi - \psi) \right] \end{pmatrix}$$



Elliptic Orbit (3)

- Torque in Inertial CS (Elliptic Approx.)

$$\begin{pmatrix} N_x \\ N_y \end{pmatrix} = -N_0 \begin{pmatrix} \sin L + \frac{e}{2} \sin(L + \ell) - \frac{3e}{2} \sin \varpi \\ -\cos L - \frac{e}{2} \cos(L + \ell) + \frac{3e}{2} \cos \varpi \end{pmatrix}$$

$$N_0 \equiv - \left[\frac{3\mu(C - A)}{2a^3} \right] \sin 2\theta (1 + 5e \cos \ell) \\ \times \left[\sin(L - \psi) + \frac{e}{2} \sin(L - \psi + \ell) - \frac{3e}{2} \sin(\varpi - \psi) \right]$$



Elliptic Orbit (4)

- Eq. of Prec. Angle and Inclination (Elliptic)

$$\frac{d\psi}{dt} = -\kappa \cos \theta \left[1 - \cos 2(L - \psi) + 3e \cos \ell - \frac{7e}{2} \cos(2L - 2\psi + \ell) + \frac{e}{2} \cos(2L - 2\psi - \ell) \right]$$

$$\frac{d\theta}{dt} = -\kappa \sin \theta \left[\sin 2(L - \psi) + \frac{7e}{2} \sin(2L - 2\psi + \ell) - \frac{e}{2} \sin(2L - 2\psi - \ell) \right]$$



Elliptic Terms

- Additional 2nd Approx. Sol. (Elliptic Terms)
 - Moon: 1 month, 10 days, 15 days

$$\delta_2^{(2)}\psi = -\cos\theta_0 \left[N_\ell \sin\ell - N_{2L+\ell} \sin(2L - 2\psi^{(1)} + \ell) + N_{2L-\ell} \sin(2L - 2\psi^{(1)} - \ell) \right]$$

$$\delta_2^{(2)}\theta = \sin\theta_0 \left[N_{2L+\ell} \cos(2L - 2\psi^{(1)} + \ell) - N_{2L-\ell} \cos(2L - 2\psi^{(1)} - \ell) \right]$$

$$N_\ell = \frac{3\kappa e}{n_\ell}, N_{2L+\ell} = \frac{7\kappa e}{2(2n_L + 2p + n_\ell)}, N_{2L-\ell} = \frac{\kappa e}{2(2n_L + 2p - n_\ell)}$$

$$N_{2L} : N_\ell \approx 1 : 6e$$

$$N_\ell : N_{2L+\ell} : N_{2L-\ell} \approx 18 : 7 : 3$$

$$n_\ell = \frac{d\ell}{dt} \approx n_L$$



Rotation Angle

- Eq. of Rot. Angle $\frac{d\phi}{dt} = \omega_c - \left(\frac{d\psi}{dt}\right) \cos \theta$
- 1st Approx.: **Unif. Rot.**

$$\phi^{(1)}(t) = \phi_0 + \omega_0 t \quad \omega_0 = \omega_{c0} + p \cos \theta_0$$

- 2nd Approx.: Variation due to **Nutation**
 - = Difference between GAST and GMST

$$\phi^{(2)}(t) = \phi_0 + \omega_0 t + \Delta\phi \quad \delta\phi = -(\Delta\psi) \cos \theta_0$$



Introduction of UT1

- UT1 = **Modified Rotation Angle of Earth**

$$UT1 \equiv \int \omega_c dt = \phi - \int \cos \theta \left(\frac{d\psi}{dt} \right) dt$$

- Assumptions: 1) Rigid Earth, 2) A=B,
3) Torque induced by External Mass points
- Important Conclusion $UT1 = \omega_c t$
 - UT1 is A Realization of Newton's Absolute Time
Independently on Perturbations



Departure point

- Newcomb's Departure Point
 - = Guinot's Non Rotating Origin (NRO)
- Definition
 - Point D on Equator (= AB Plane) of Principal CS such that UT1 = Arc Length of DA

$$UT1 = \widehat{DA} = \widehat{DN} + \widehat{NA}$$

$$\widehat{NA} = \phi, \quad \widehat{DN} = \int \cos \theta d\psi$$

Inclined Orbit

- Pert. Body in Inertial CS
 - Origin=Rotating Body's Barycenter
- Coordinates Expression

$$\mathbf{r} = \begin{pmatrix} r_X \\ r_Y \\ r_Z \end{pmatrix}$$

$$r_C = \mathbf{r} \cdot \mathbf{e}_C = r_X \sin \theta \sin \psi - r_Y \sin \theta \cos \psi + r_Z \cos \theta$$

$$\mathbf{r} \times \mathbf{e}_C = \begin{pmatrix} r_Y \cos \theta + r_Z \sin \theta \cos \psi \\ -r_X \cos \theta + r_Z \sin \theta \sin \psi \\ -r_X \sin \theta \cos \psi - r_Y \sin \theta \sin \psi \end{pmatrix}$$



Inclined Orbit (2)

- Torque in Inertial CS (Inclined Orbit)

$$N_x = \frac{-3\mu(C-A)}{r^5} \left[r_x r_y \sin \theta \cos \theta \sin \psi + r_x r_z \sin^2 \theta \sin \psi \cos \psi \right. \\ \left. + r_y r_z (\cos^2 \theta - \sin^2 \theta \cos^2 \psi) + (r_z^2 - r_y^2) \sin \theta \cos \theta \cos \psi \right]$$

$$N_y = \frac{-3\mu(C-A)}{r^5} \left[r_x r_y \sin \theta \cos \theta \cos \psi - r_y r_z \sin^2 \theta \sin \psi \cos \psi \right. \\ \left. - r_x r_z (\cos^2 \theta - \sin^2 \theta \cos^2 \psi) + (r_z^2 - r_x^2) \sin \theta \cos \theta \sin \psi \right]$$



Inclined Orbit (3)

- Eq. of Precession/Inclination (Poisson Approx.)

- Q: Show

$$\frac{d\psi}{dt} = \frac{3\mu(C-A)}{C\omega_c r^5} \left[r_x r_y \cos \theta \sin 2\psi + r_y r_z \left(\frac{\cos 2\theta}{\sin \theta} \right) \cos \psi \right. \\ \left. - r_x r_z \left(\frac{\cos 2\theta}{\sin \theta} \right) \sin \psi - r_x^2 \cos \theta \sin^2 \psi - r_y^2 \cos \theta \cos^2 \psi + r_z^2 \cos \theta \right]$$

$$\frac{d\theta}{dt} = \frac{3\mu(C-A)}{C\omega_c r^5} \left[-r_x r_y \sin \theta \cos 2\psi - r_y r_z \cos \theta \sin \psi \right. \\ \left. + r_x r_z \cos \theta \cos \psi + (r_x^2 - r_y^2) \sin \theta \sin \psi \cos \psi \right]$$

Inclined Circular Orbit

- **Inclined** but Circular Orbit $r = \bar{r}$
 - Ex.: Moon in Earth Rotation
 - Constant Orbital Inclination $I = I_0$
 - **Slowly Moving Ascending Node** $\Omega = \Omega_0 + n_\Omega t$
 - Argument of Latitude $F \equiv L - \Omega$
- Pert. Body's Position in Inertial CS (Q: Show)

$$\mathbf{r} = \begin{pmatrix} r_X \\ r_Y \\ r_Z \end{pmatrix} = \bar{r} \begin{pmatrix} \cos F \cos \Omega - \sin F \cos I \sin \Omega \\ \cos F \sin \Omega + \sin F \cos I \cos \Omega \\ \sin F \sin I \end{pmatrix}$$

Corrected Precession

- Approx. 1: Zero Prec. Angle, Const. Inclination
- Approx. 2: Removing Short-Period Terms by **Averaging Method**

$$\left\langle \frac{d\psi}{dt} \right\rangle = \kappa \cos \theta_0 \left[-\left(1 - \frac{3}{2} \sin^2 I_0\right) + \frac{\sin 2I_0}{\tan 2\theta_0} \cos \Omega + \frac{\sin^2 I_0}{2} \cos 2\Omega \right]$$

$$\left\langle \frac{d\theta}{dt} \right\rangle = -\kappa \cos \theta_0 \left[\frac{\sin 2I_0}{2} \sin \Omega + \frac{\tan \theta_0 \sin^2 I_0}{2} \sin 2\Omega \right]$$

- Prec. Const. **Correction**

$$p = \kappa \cos \theta_0 \left(1 - \frac{3}{2} \sin^2 I_0\right)$$

Averaging Method

- = Removal of Short-Periodic Terms
- Meaning: Medium-scale Time Average
- Step 1: Fourier Expansion w.r.t. Fast Angles
- Step 2: Trig. F. of Fast Angles = 0

$$\langle \sin jF \rangle = 0 \quad \langle \cos jF \rangle = 0 \quad \text{when } j \neq 0$$

- Ex.: F = Fast Angle

$$\langle r_z^2 \cos \theta \rangle = \langle \bar{r}^2 \sin^2 F \sin^2 I \cos \theta \rangle = \frac{\bar{r}^2 \sin^2 I_0 \cos \theta_0}{2}$$

Long Period Nutation

- Assuming Const. Speed of Ascending Node
 - Note: **Clockwise** Motion $n_{\Omega} < 0$

$$\Delta^{(1)}\psi = \frac{\kappa \cos \theta_0}{n_{\Omega}} \left[\frac{\sin 2I_0}{\tan 2\theta_0} \sin \Omega + \frac{\sin^2 I_0}{4} \sin 2\Omega \right]$$

$$\Delta^{(1)}\theta = \frac{\kappa \cos \theta_0}{n_{\Omega}} \left[\frac{\sin 2I_0}{2} \cos \Omega + \frac{\tan \theta_0 \sin^2 I_0}{4} \cos 2\Omega \right]$$

- Long Period = Large Amplitude
 - So-called **Nutation**: 18.6y



Luni-solar Precession (2)

- Note: **Negative** $\psi \equiv -\psi^{(1)} = pt$, $\varepsilon \equiv -\theta^{(1)} = \varepsilon_0$
- Moon (without prime), Sun (with prime)
- Obliquity $\varepsilon_0 \approx 84381''$, $\cos \varepsilon_0 \approx 0.917482$, $\sin \varepsilon_0 \approx 0.397777$
- Orbit Inclination $I_0 \approx 18467''$, $\sin I_0 \approx 0.089411$
- Estimate of Luni-solar Precession

$$\kappa \cos \varepsilon_0 \left(1 - \frac{3}{2} \sin^2 I_0 \right) \approx 33.7418''/\text{jy}$$

$$p \approx 49.7372''/\text{jy}$$

$$\kappa' \cos \varepsilon_0 \approx 15.9954''/\text{jy}$$

Long P. Nutation (2)

- Note: **Negative** $\Delta\psi = -\Delta^{(1)}\psi$, $\Delta\varepsilon = -\Delta^{(1)}\theta$
- Moon's Effect Only

$$\Delta\psi = \frac{\kappa \cos \varepsilon_0}{|n_\Omega|} \left[-\left(\frac{\sin 2I}{\tan 2\varepsilon_0} \right) \sin \Omega + \frac{\sin^2 I}{4} \sin 2\Omega \right]$$

$$\Delta\varepsilon = \frac{\kappa \cos \varepsilon_0}{|n_\Omega|} \left[\frac{\sin 2I}{2} \cos \Omega - \frac{\tan \varepsilon_0 \sin^2 I}{4} \cos 2\Omega \right]$$

Long P. Nutation (3)

- Speed of Nodal Motion $\frac{\kappa \cos \varepsilon_0}{|n_\Omega|} \approx 103.130''$
 $n_\Omega \approx -6.96289 \times 10^4''/\text{jy}$

- Amplitude of LP Nutation (Q: Confirm)

$$\psi_\Omega = \frac{-\kappa \cos \varepsilon_0}{|n_\Omega|} \left(\frac{\sin 2I}{\tan 2\varepsilon_0} \right) = -17.202''$$

$$\varepsilon_\Omega = \frac{\kappa \cos \varepsilon_0}{|n_\Omega|} \left(\frac{\sin 2I}{2} \right) = 9.184''$$

$$\psi_{2\Omega} = \frac{\kappa \cos \varepsilon_0}{|n_\Omega|} \left(\frac{\sin^2 I}{4} \right) = 0.206''$$

$$\varepsilon_{2\Omega} = \frac{-\kappa \cos \varepsilon_0}{|n_\Omega|} \left(\frac{\tan \varepsilon_0 \sin^2 I}{4} \right) = -0.089''$$

- Summary $\begin{cases} \Delta\psi = -17.202'' \sin \Omega + 0.206'' \sin 2\Omega \\ \Delta\varepsilon = +9.184'' \sin \Omega - 0.089'' \sin 2\Omega \end{cases}$



Short Period Nutation

- Simplification of Expression
 - $L \rightarrow L + \psi$: Adding Precession A. To Mean Long.
 - = Measuring Longitude from **Moving** Equinox

$$\delta\psi = \cos \varepsilon_0 \left[-N_{2L} \sin 2L + N_{\ell} \sin \ell - N_{2L+\ell} \sin(2L + \ell) + N_{2L-\ell} \sin(2L - \ell) \right. \\ \left. - N_{2L'} \sin 2L' + N_{\ell'} \sin \ell' - N_{2L'+\ell'} \sin(2L' + \ell') + N_{2L'-\ell'} \sin(2L' - \ell') \right]$$

$$\delta\varepsilon = \sin \varepsilon_0 \left[N_{2L} \cos 2L + N_{2L+\ell} \cos(2L + \ell) - N_{2L-\ell} \cos(2L - \ell) \right. \\ \left. + N_{2L'} \cos 2L' + N_{2L'+\ell'} \cos(2L' + \ell') - N_{2L'-\ell'} \cos(2L' - \ell') \right]$$



Short P. Nutation (2)

- Constants (Moon=Unprimed, Sun=Primed)

$$n_\ell \approx 1.718 \times 10^7 \text{ "/y} \quad n_{\ell'} \approx 1.296 \times 10^6 \text{ "/y} \quad e_M \approx 0.0549, e_E \approx 0.0167$$

- Amplitude Coefficient $N_{2L} = \frac{\kappa}{2n_\ell}, N_\ell = \frac{3\kappa e}{n_\ell}, N_{2L+\ell} = \frac{7\kappa e}{6n_\ell}, N_{2L-\ell} = \frac{\kappa e}{2n_\ell}$

- Numerical Evaluation (Q: Confirm)

$$N_{2L} \approx 0.228", N_\ell \approx 0.075", N_{2L+\ell} \approx 0.029", N_{2L-\ell} \approx 0.013"$$

$$N_{2L'} \approx 1.383", N_{\ell'} \approx 0.139", N_{2L'+\ell'} \approx 0.054", N_{2L'-\ell'} \approx 0.023"$$



Short P. Nutation (3)

- Summary (Q: Compare with Latest Theory)

$$\begin{aligned}\delta\psi = & -1.269'' \sin 2L' - 0.209'' \sin 2L + 0.128'' \sin \ell' \\ & + 0.069'' \sin \ell - 0.050'' \sin (2L' + \ell') - 0.027'' \sin (2L + \ell) \\ & + 0.021'' \sin (2L' - \ell') + 0.012'' \sin (2L - \ell)\end{aligned}$$

$$\begin{aligned}\delta\varepsilon = & 0.550'' \cos 2L' + 0.091'' \cos 2L + 0.021'' \cos (2L' + \ell') \\ & + 0.012'' \cos (2L + \ell) - 0.009'' \cos (2L' - \ell') \\ & - 0.005'' \cos (2L - \ell)\end{aligned}$$

- Analytical Dynamics is Needed for Detailed Eval.

$A < B < C$

- Analytical Treatment is Quite Difficult
- **Post-Poisson Approx.** $\omega_A, \omega_B \ll \omega_C$
 - Eq. of Rotation for Tidal Torque

$$\frac{d}{dt} \begin{pmatrix} \omega_A \\ \omega_B \\ \omega_C \end{pmatrix} + \begin{pmatrix} \alpha \omega_B \omega_C \\ \beta \omega_C \omega_A \\ 0 \end{pmatrix} = \frac{3\mu}{r^5} \begin{pmatrix} \alpha r_B r_C \\ \beta r_C r_A \\ \gamma r_A r_B \end{pmatrix}$$

Cassini Approximation

- Pert. Body: Inclined Circular Orbit Again
- **Cassini's Laws** Hold (Ex.: Moon)

$$\langle \psi \rangle = \Omega, \quad \langle \theta \rangle = -\theta_0, \quad \langle \phi + \psi \rangle = L + \pi$$

- Assuming Small Inclination: I and θ

- Approx. Expr.

(Q: Show)

$$F = L - \Omega$$

$$\begin{pmatrix} r_A \\ r_B \\ r_C \end{pmatrix} = \bar{r} \begin{pmatrix} 1 \\ 0 \\ -(\theta_0 + I) \sin F \end{pmatrix}$$

Cassini Approx. (2)

- Approximated Eq. of Rotation

$$\frac{d}{dt} \begin{pmatrix} \omega_A \\ \omega_B \\ \omega_C \end{pmatrix} + \omega_C \begin{pmatrix} \alpha \omega_B \\ \beta \omega_A \\ 0 \end{pmatrix} = \frac{-3\mu\beta(\theta_0 + I)}{\bar{r}^3} \begin{pmatrix} 0 \\ \sin F \\ 0 \end{pmatrix}$$

- An Integral Exists

- C-axis Comp. of Angular Velocity
= Mean Orbital Longitude Motion

$$\omega_C = n_L$$

Cassini Approx. (3)

- A- & B-Comp.: Eq. of **Forced** Oscillation

$$\frac{d}{dt} \begin{pmatrix} \omega_A \\ \omega_B \end{pmatrix} + n_L \begin{pmatrix} \alpha \omega_B \\ \beta \omega_A \end{pmatrix} = W \begin{pmatrix} 0 \\ \sin F \end{pmatrix} \quad W \equiv \frac{-3\mu\beta(\theta_0 + I)}{\bar{r}^3}$$

- Special Solution (Q: Confirm)

$$\begin{pmatrix} \omega_A \\ \omega_B \end{pmatrix} = \frac{W}{n_F^2 - n_E^2} \begin{pmatrix} \alpha n_L \sin F \\ -n_F \cos F \end{pmatrix} \quad \begin{matrix} n_E \equiv n_L \sqrt{-\alpha\beta} \ll n_L \\ n_F \equiv \frac{dF}{dt} = n_L - n_\Omega > n_L \end{matrix}$$

- Q: Obtain General Sol. including Free Osc.

Cassini Approx. (4)

- Approx. Solution of Euler Angles

$$\psi = \psi_0 + \frac{W}{2\theta_0(n_F^2 - n_E^2)} \left[(n_F - \alpha n_L)(t - t_0) - \left(\frac{n_F + \alpha n_L}{2n_F} \right) \sin 2F \right]$$

$$\theta = -\theta_0 + \frac{W}{2(n_F^2 - n_E^2)} \left(\frac{n_F + \alpha n_L}{2n_F} \right) \cos 2F$$

$$\phi = L + \pi - \psi$$

- Satisfying Two Consistency Conditions

$$\langle \theta \rangle = -\theta_0, \quad \langle \phi + \psi \rangle = L + \pi$$

Cassini Approx. (5)

- Last Consistency Condition $\langle \psi \rangle = \Omega$

- Determination of Mean Inclination

$$\left\langle \frac{d\psi}{dt} \right\rangle = n_{\Omega} \rightarrow \theta_0 = \frac{n_I}{(-n_{\Omega}) - n_I} I \quad n_I \equiv \frac{-3\mu\beta(n_F - \alpha n_L)}{2\bar{r}^3(n_F^2 - n_E^2)}$$

- Moon: (Q: Compare with Eckhardt(1981))

$$n_{\Omega} \approx -6.963 \times 10^4 \text{"/jy}, \quad n_F \approx 1.7395 \times 10^7 \text{"/jy}, \quad n_L \approx 1.7325 \times 10^7 \text{"/jy}$$

$$\alpha_M = 3.99 \times 10^{-4}, \quad \beta_M = -6.32 \times 10^{-4}, \quad n_E = 8.70 \times 10^{-3}$$

$$\mu/\bar{r}^3 \approx n_F^2 \quad n_I \approx 1.648 \times 10^4 \text{"/jy} \quad I \approx 18467'' \quad \therefore \theta_0 \approx 5527''$$

Post-Cassini Approx.

- **Extension** of Cassini Approximation

$$\langle \psi \rangle = \Omega, \langle \theta \rangle = -\theta_0, \langle \phi \rangle = \langle v \rangle + \pi$$

- Argument of Latitude $v = \omega + f$, True Anomaly f
- Assuming Small Orb. Incl. I , and Incl. θ
- Circular Orbit

- Approx. Expr.
(Q: Show)

$$\begin{pmatrix} r_A \\ r_B \\ r_C \end{pmatrix} \approx \bar{r} \begin{pmatrix} 1 \\ \psi + \phi - (v + \Omega + \pi) \\ -(I - \theta) \sin v \end{pmatrix}$$

Post-Cassini Approx. (2)

- Approximate Equation of Rotation

$$\frac{d}{dt} \begin{pmatrix} \omega_A \\ \omega_B \\ \omega_C \end{pmatrix} + \omega_C \begin{pmatrix} \alpha \omega_B \\ \beta \omega_A \\ 0 \end{pmatrix} = \frac{3\mu}{r^3} \begin{pmatrix} 0 \\ \beta(I - \theta) \sin v \\ \gamma(v + \Omega + \pi - \psi - \phi) \end{pmatrix}$$

- $\theta \sim 0 \rightarrow$ C-axis Comp.: **Forced Oscillation**

$$\frac{d^2(\psi + \phi)}{dt^2} + n_P^2(\psi + \phi) = n_P^2(v + \Omega + \pi) \quad n_P \equiv \sqrt{\frac{3\gamma\mu}{r^3}}$$

Post-Cassini Approx. (3)

- Libration in Longitude (=Diff. from L)
 - **Optical** (=Apparent) Libr. $\nu + \Omega - L$
 - **Physical** (=Real) Libr. $\tau \equiv \psi + \phi - L - \pi$
- Eq. of Forced Oscillation (Q: Show)
 - Angular Velocity of Free Oscillation: n_p
 - RHS is derived from Orbital Theory

$$\frac{d^2\tau}{dt^2} + n_p^2\tau = n_p^2(\nu + \Omega - L) = n_p^2 \sum_k H_k \sin(n_k t + \phi_k)$$

Post-Cassini Approx. (4)

- Solution of **Physical** Libration in Long.
 - Suffix P = Physical
 - Q: Derive

$$\tau = \tau_0 + \tau_P \sin n_P (t - t_0) + \sum_k \frac{n_P^2}{n_P^2 - n_k^2} H_k \sin(n_k t + \varphi_k)$$

- Solution of C-axis Comp. of Ang. Vel.

$$\omega_C = n_L + n_P \tau_P \cos n_P (t - t_0) + \sum_k \frac{n_k n_P^2}{n_P^2 - n_k^2} H_k \cos(n_k t + \varphi_k)$$

Post-Cassini Approx. (5)

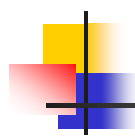
- A-, B-Axis Comp.: $\theta \sim -\theta_0 \rightarrow$ **Forced Osc.**

$$\frac{d}{dt} \begin{pmatrix} \omega_A \\ \omega_B \end{pmatrix} + n_L \begin{pmatrix} \alpha \omega_B \\ \beta \omega_A \end{pmatrix} = \begin{pmatrix} 0 \\ \sum_k J_k \sin(n_k t + \varphi_k) \end{pmatrix}$$

- General Solution (Q: Confirm)

- Ang. Vel. of Free Polar Motion $n_E \equiv n_L \sqrt{-\alpha\beta}$

$$\begin{pmatrix} \omega_A \\ \omega_B \end{pmatrix} = \omega_0 \begin{pmatrix} \sqrt{\alpha} \cos n_E (t - t_0) \\ \sqrt{-\beta} \sin n_E (t - t_0) \end{pmatrix} + \sum_k \frac{J_k}{n_k^2 - n_E^2} \begin{pmatrix} \alpha n_L \sin(n_k t + \varphi_k) \\ -n_k \cos(n_k t + \varphi_k) \end{pmatrix}$$



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